



***Probing the Early Evolution of Massive  
Black Holes with Dwarf Starburst Galaxies***

**Amy Reines**

Einstein Fellow

National Radio Astronomy Observatory



# Supermassive black holes and galaxy evolution

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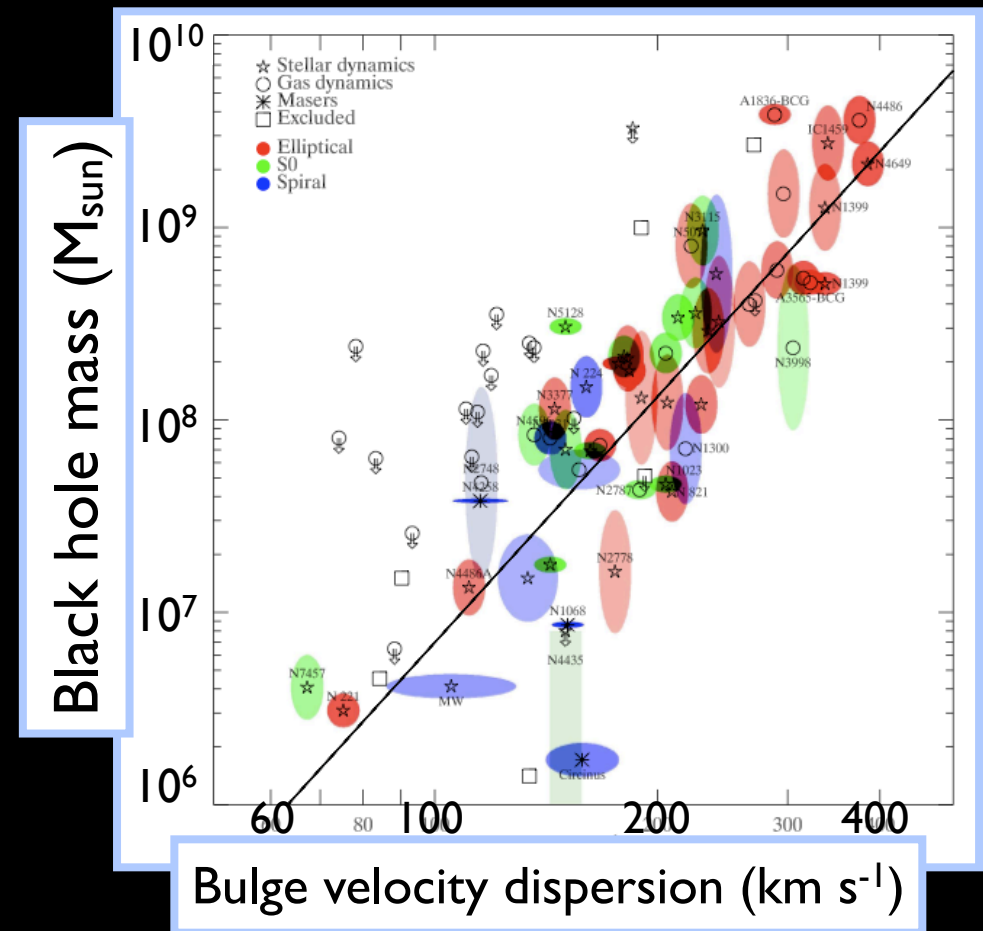


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## Some questions:

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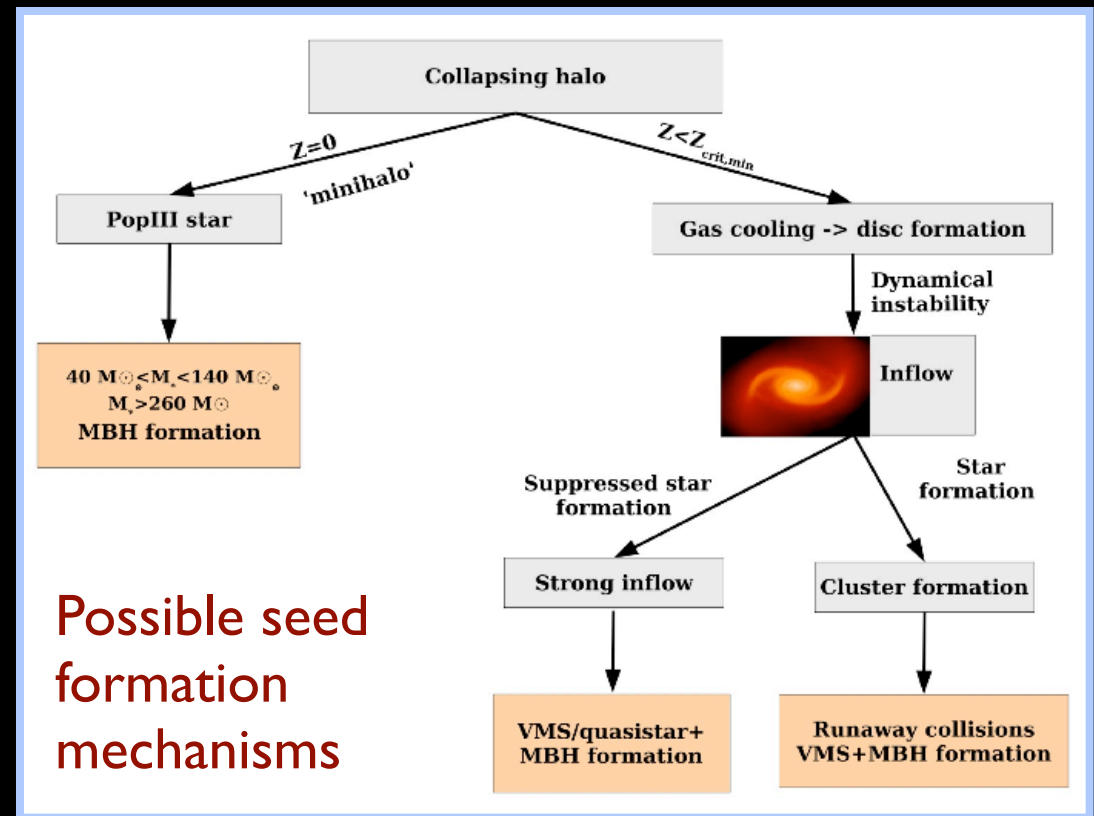


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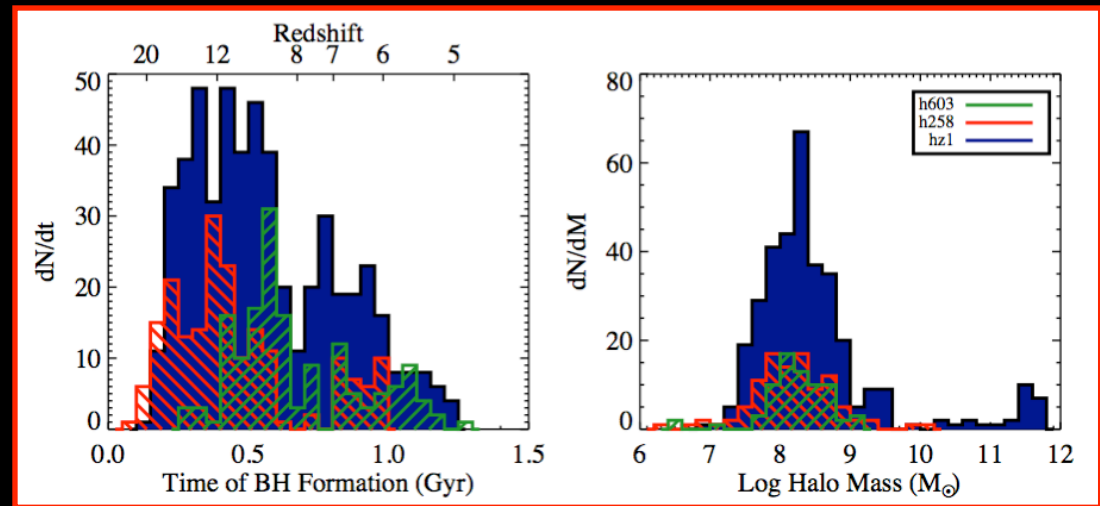


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Bellovary et al. (2011)

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- Did galaxies and nuclear black holes grow synchronously? If not, which developed first?
- How do the “seeds” of supermassive black holes form?
- When did the seeds form and what types of galaxies did they form in?
- How common are massive black holes in modern bulgeless and dwarf galaxies?

Pox 52



NGC 4395





*“An actively accreting massive black hole in the dwarf starburst galaxy Henize 2-10”*

Reines, Sivakoff, Johnson & Brogan 2011, Nature, 470, 66



*“Astrophysics: Big black hole found in tiny galaxy”*

Greene 2011, Nature, 470, 45

## Henize 2-10

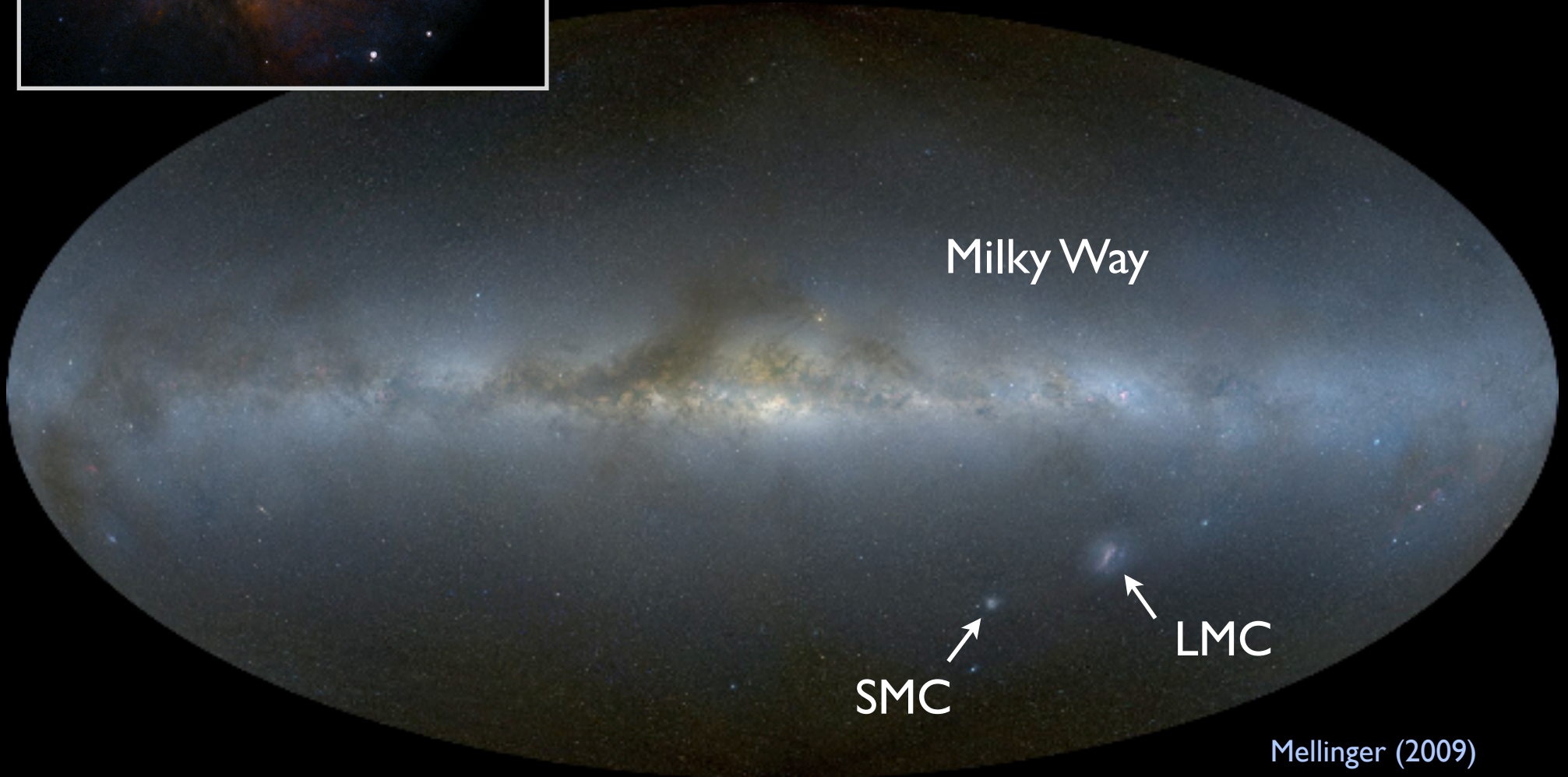


- Nearby ( $D \sim 9$  Mpc) dwarf starburst galaxy (Allen et al. 1976)
- Compact ( $\sim 1$  kpc), irregular morphology
- Young super star clusters (proto-globular clusters) (e.g. Johnson et al. 2000)



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- Compact ( $\sim 1$  kpc), irregular morphology
- Young super star clusters (proto-globular clusters) (e.g. Johnson et al. 2000)
- Main optical body is about half the size of the SMC
- SFR  $\sim 10$  times the LMC but similar stellar and HI masses





# Observations

Infant super star clusters:

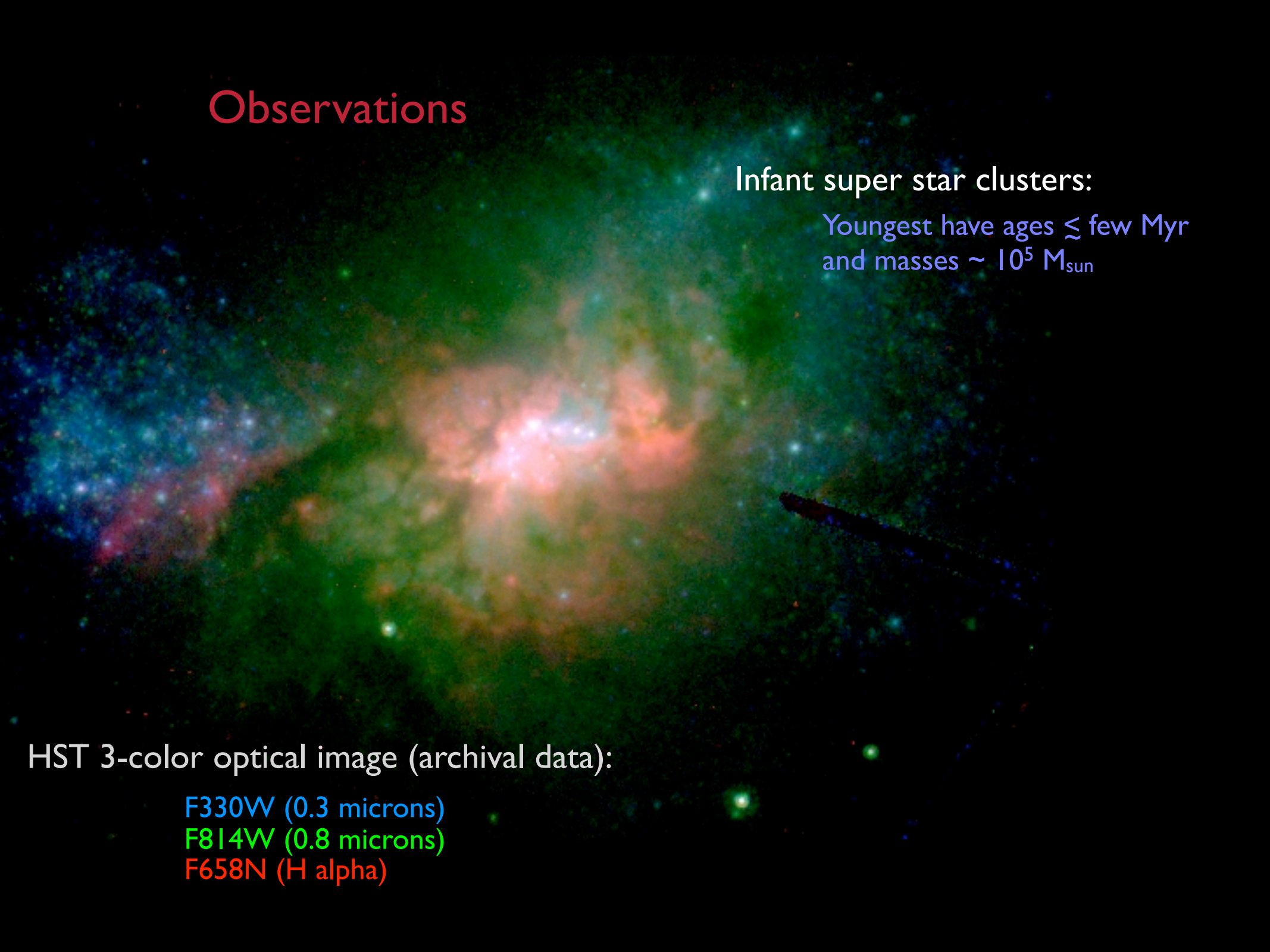
Youngest have ages  $\lesssim$  few Myr  
and masses  $\sim 10^5 M_{\text{sun}}$

HST 3-color optical image (archival data):

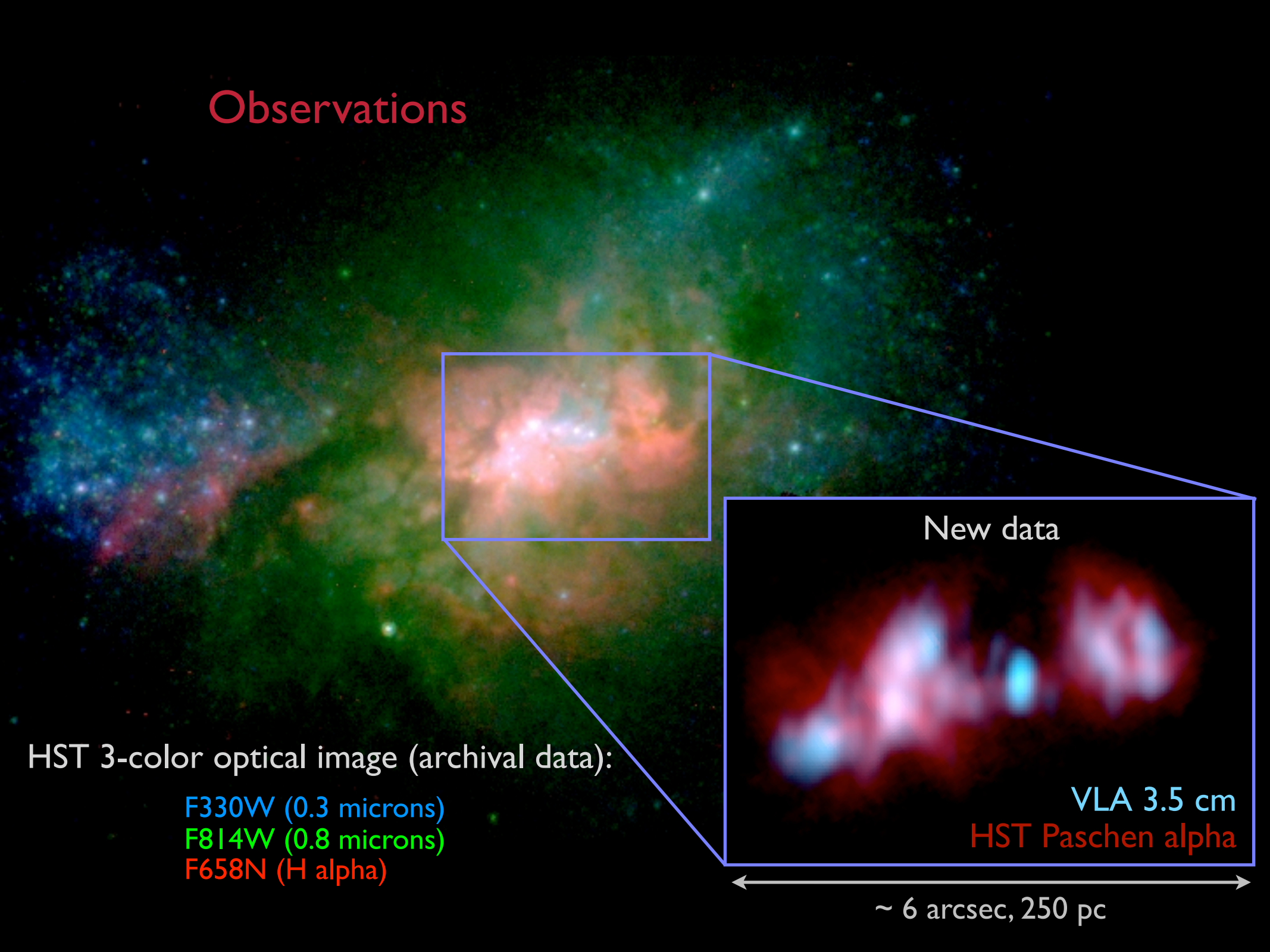
F330W (0.3 microns)

F814W (0.8 microns)

F658N (H alpha)



# Observations



HST 3-color optical image (archival data):

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New data

VLA 3.5 cm

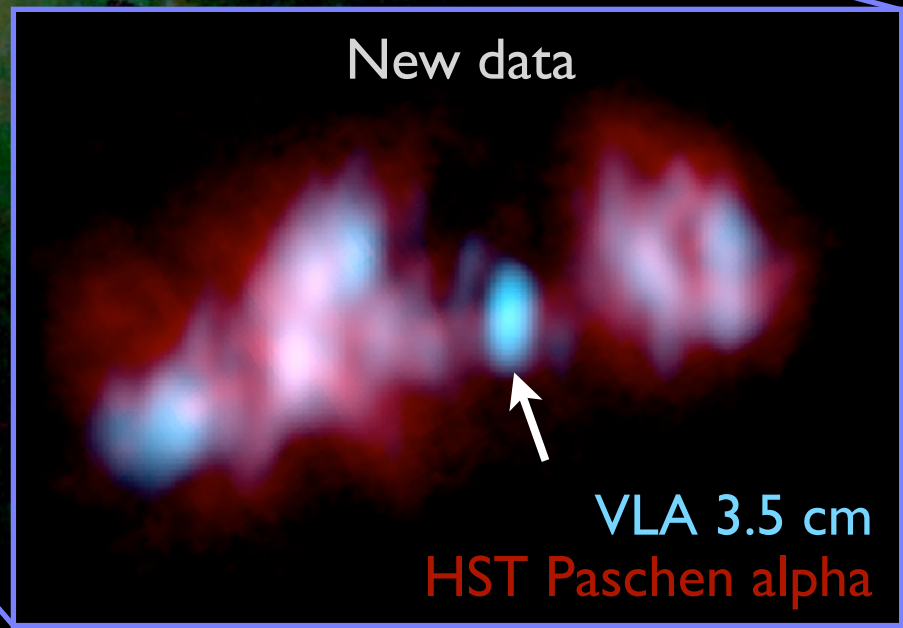
HST Paschen alpha

~ 6 arcsec, 250 pc



# Observations

New focus:  
the central source



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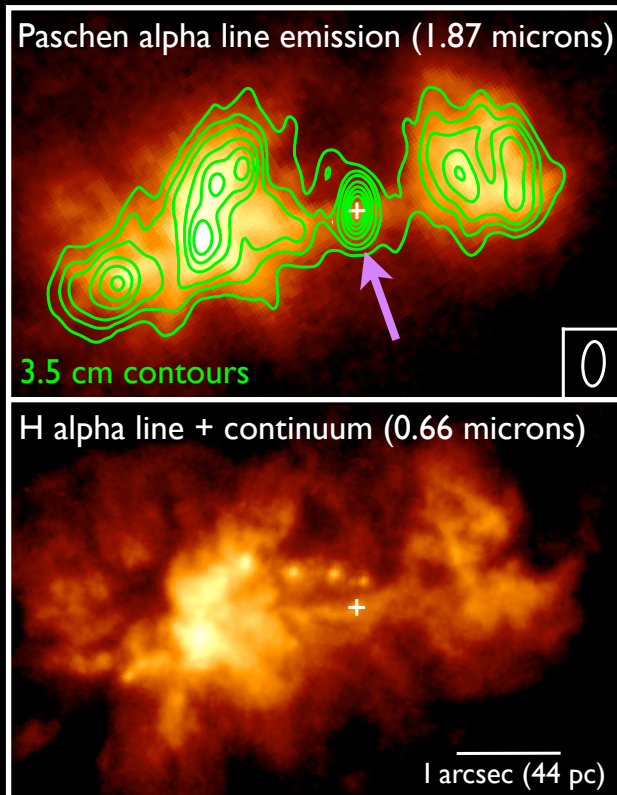
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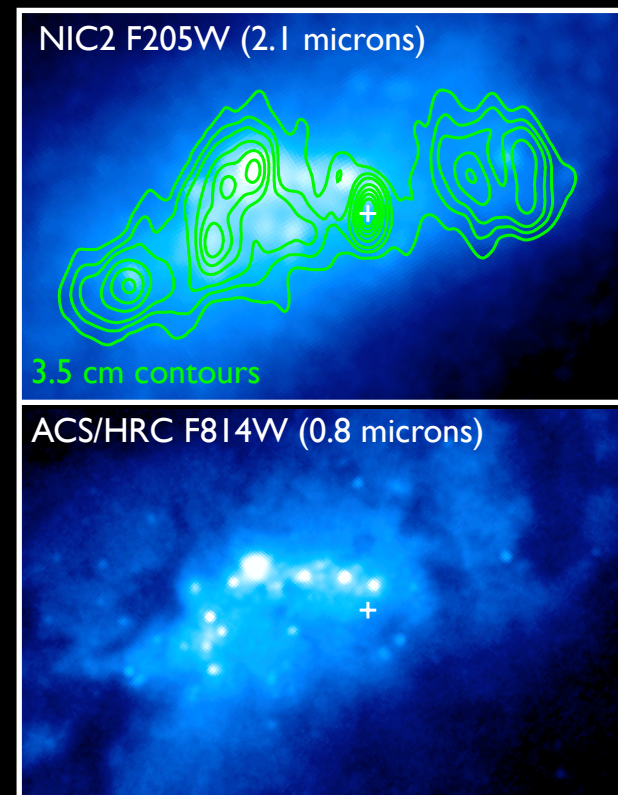


# The central source in Henize 2-10

## Narrow-band imaging (ionized gas)

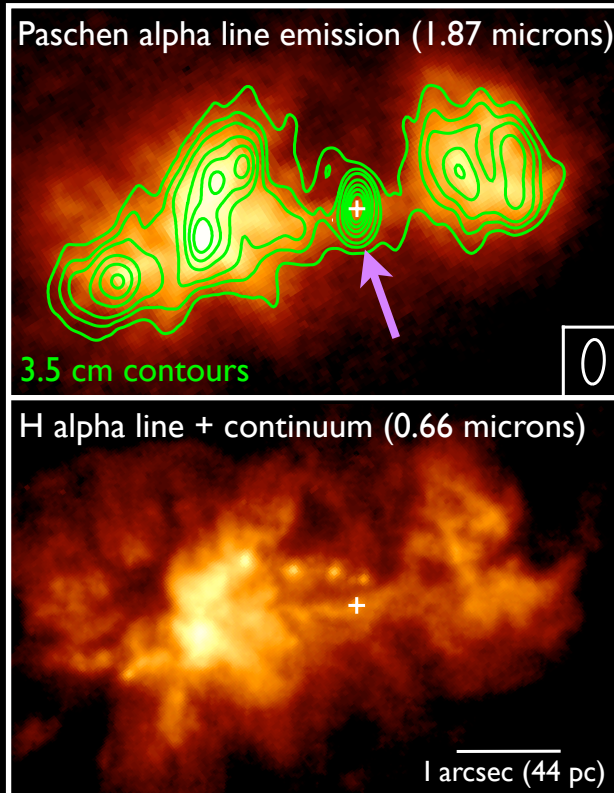


## Broad-band imaging (stars)

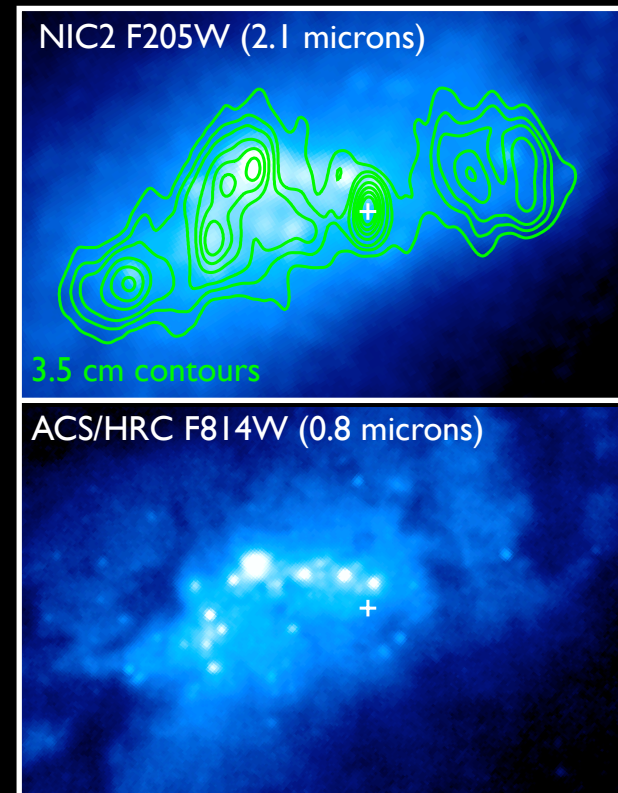


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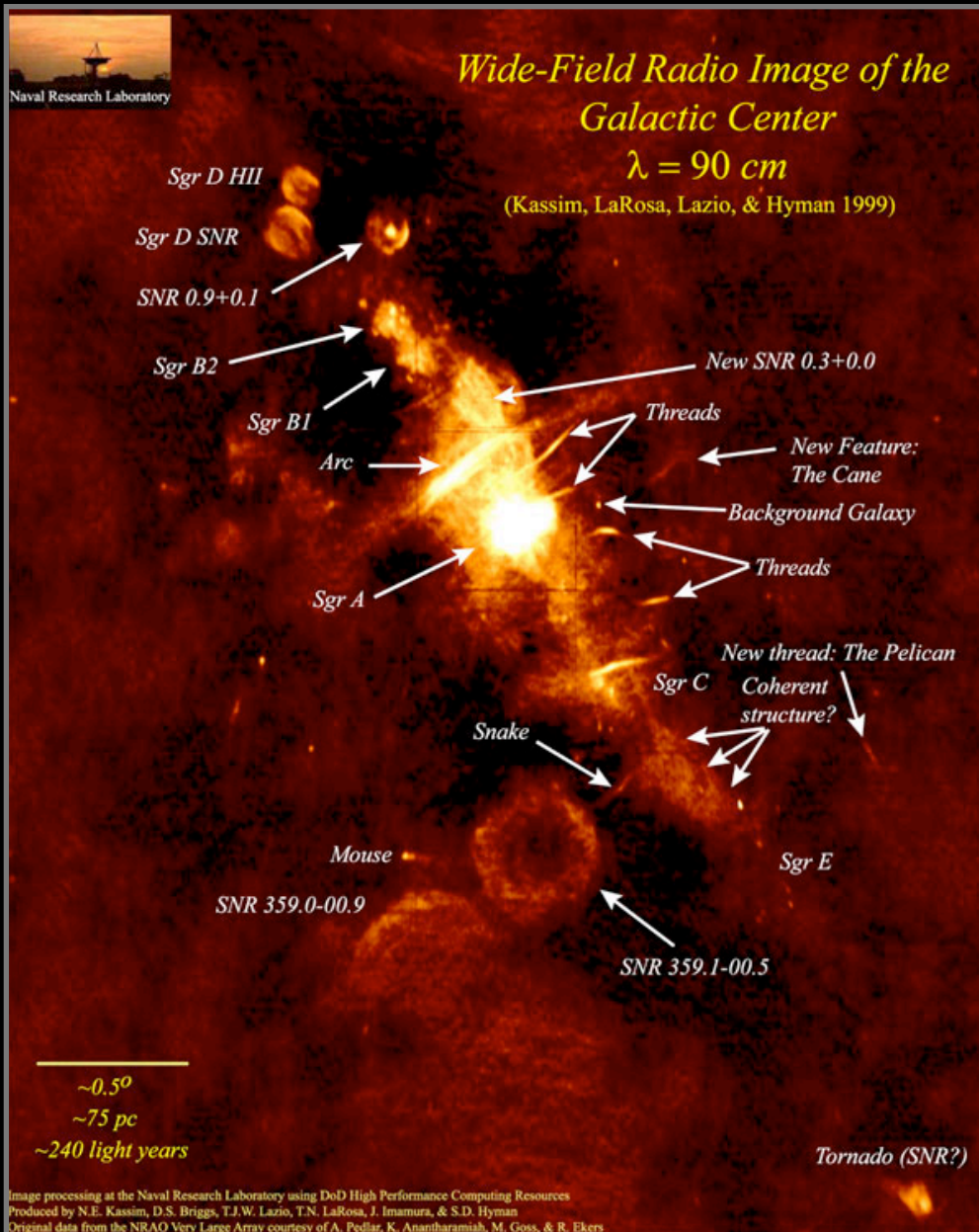


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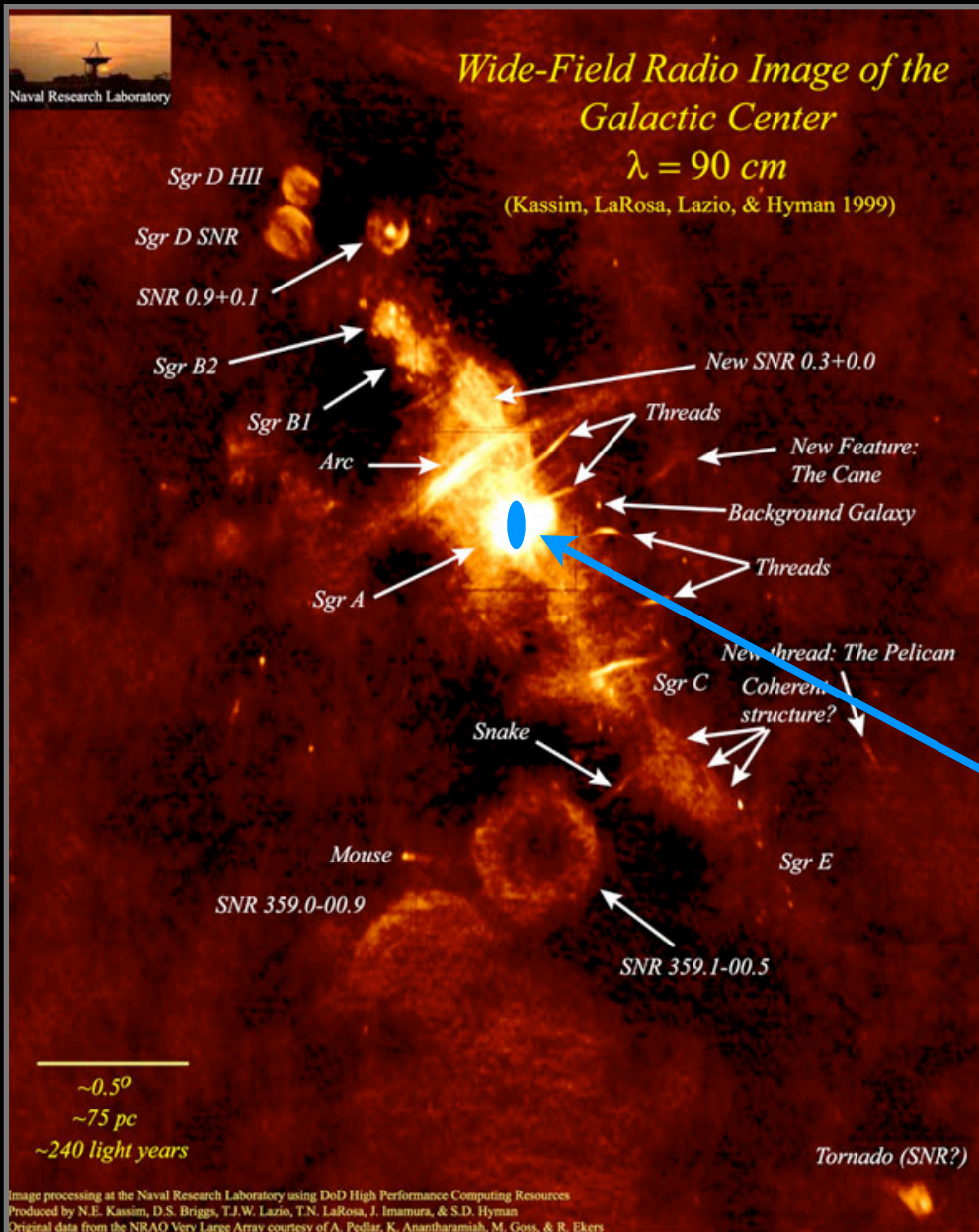
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Radio image of the Galactic Center



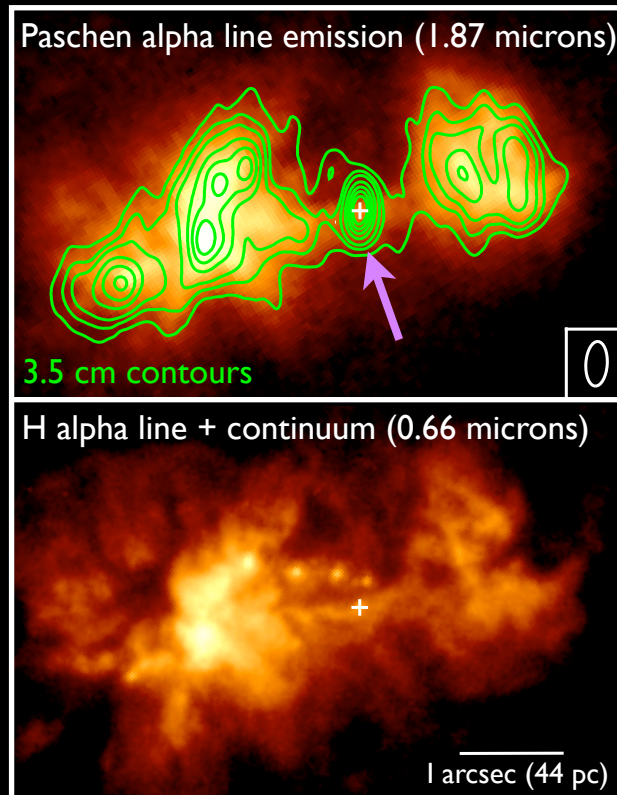
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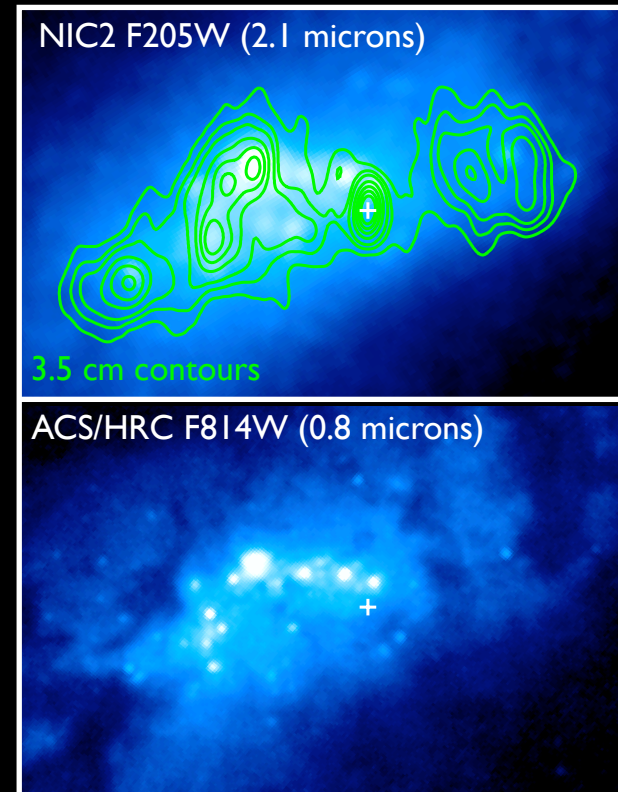
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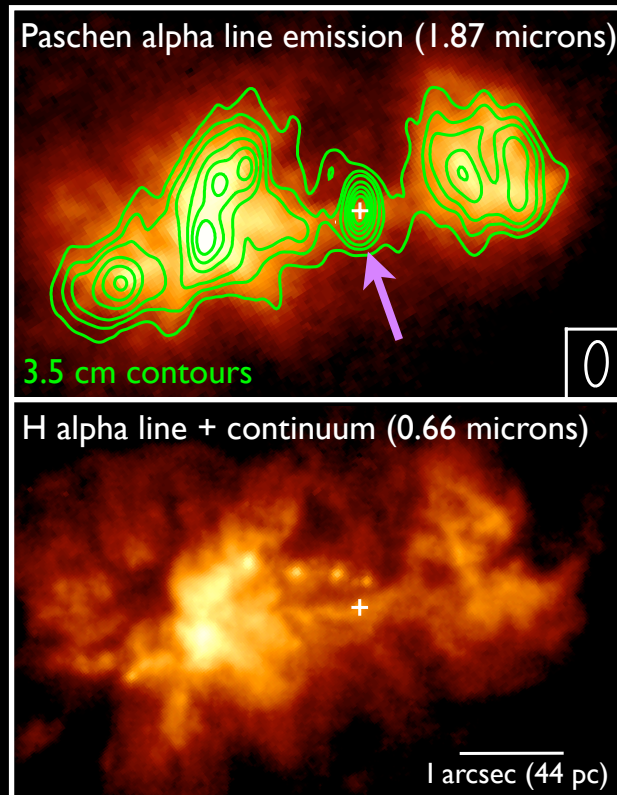
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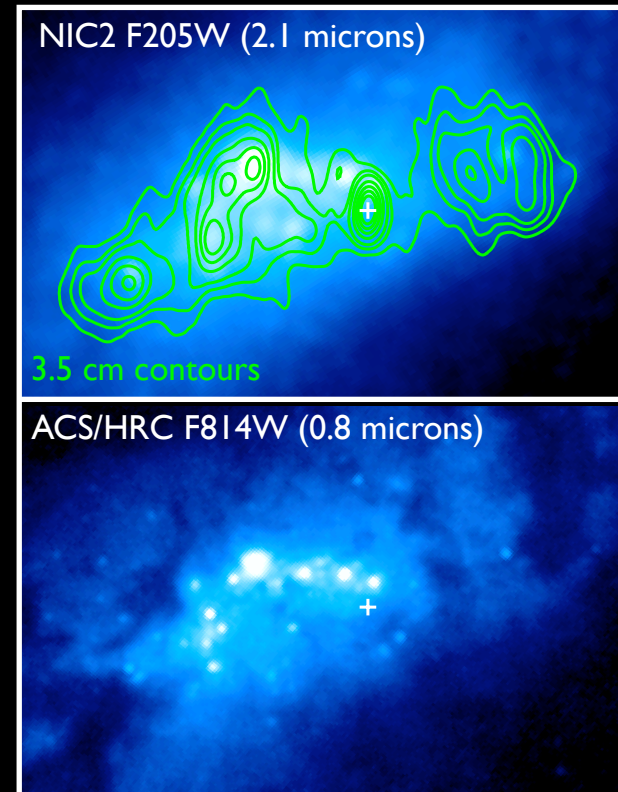
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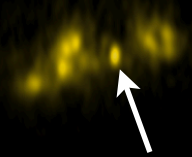
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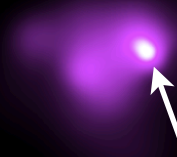
Hubble Space Telescope



Very Large Array radio telescope



Chandra X-ray Observatory





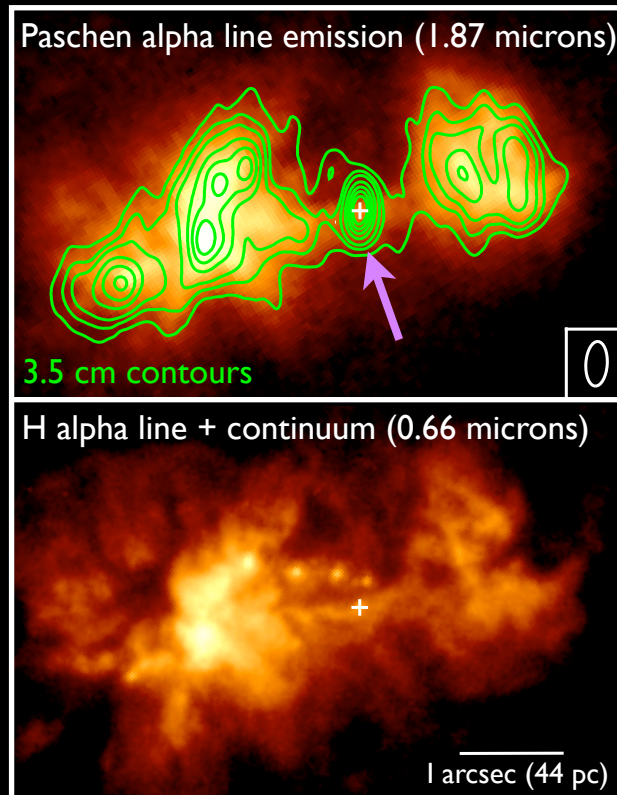
Central region strongly emitting  
radio waves and energetic X-rays



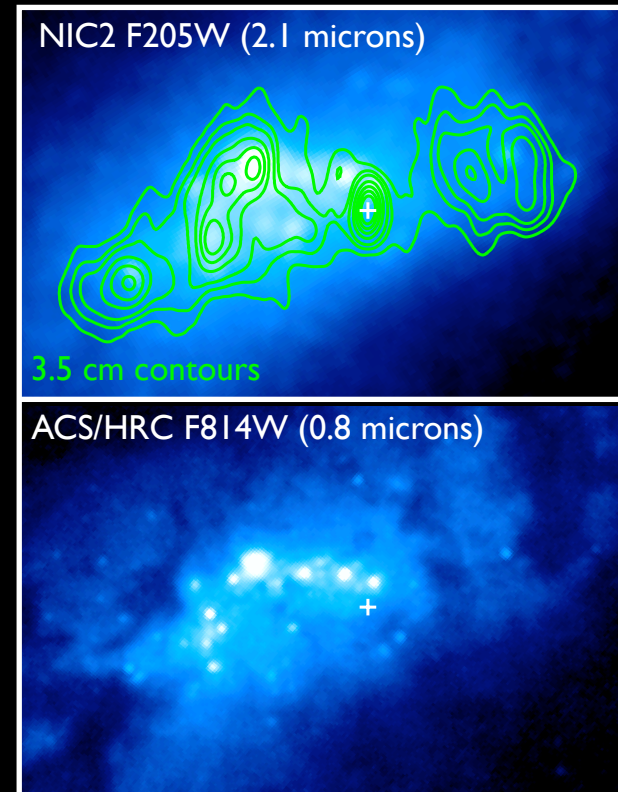
Images from <http://chandra.harvard.edu/press>

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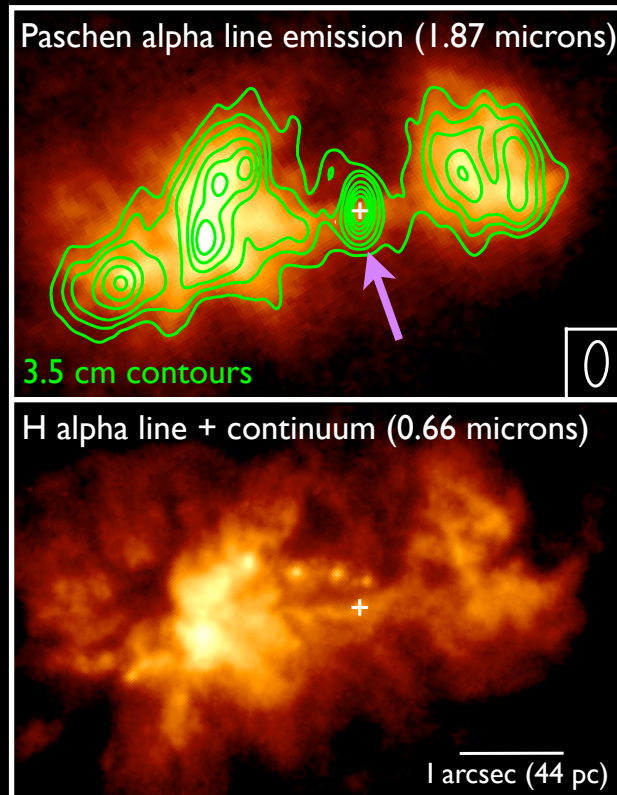


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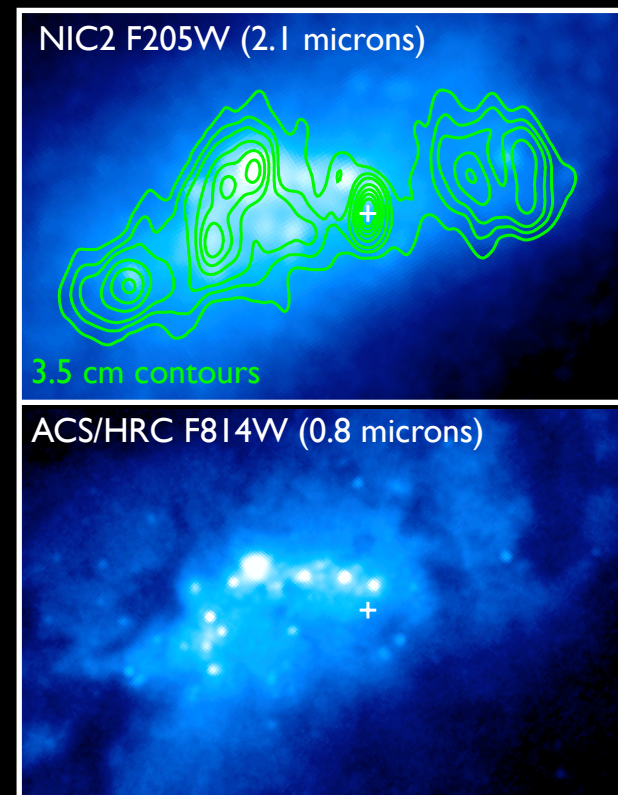


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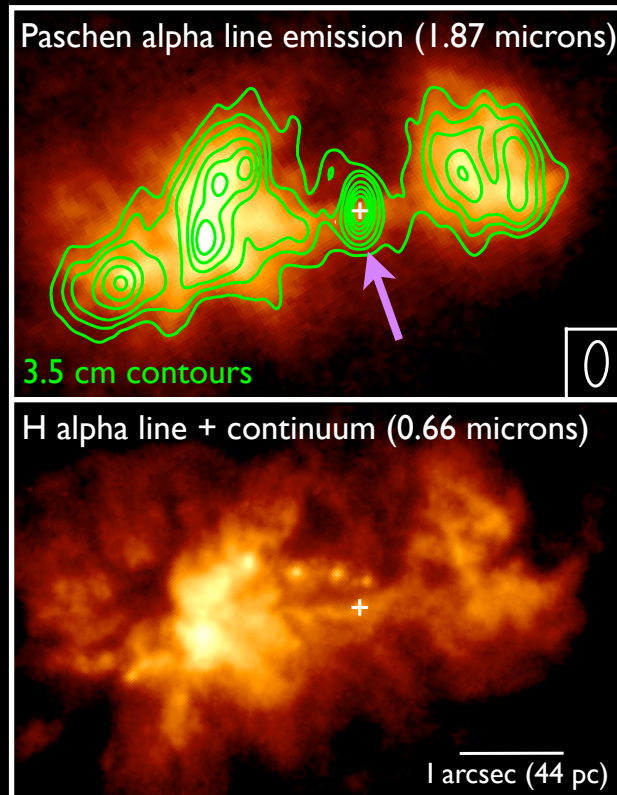


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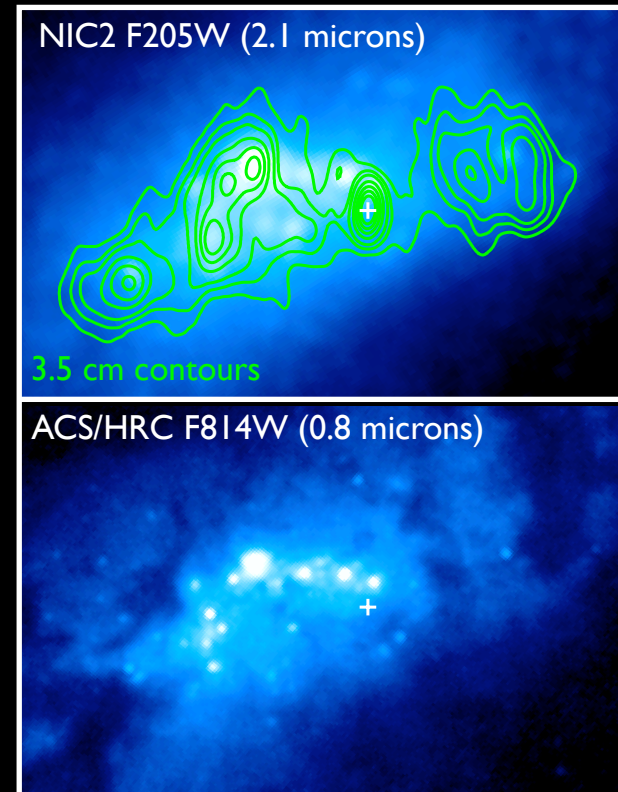


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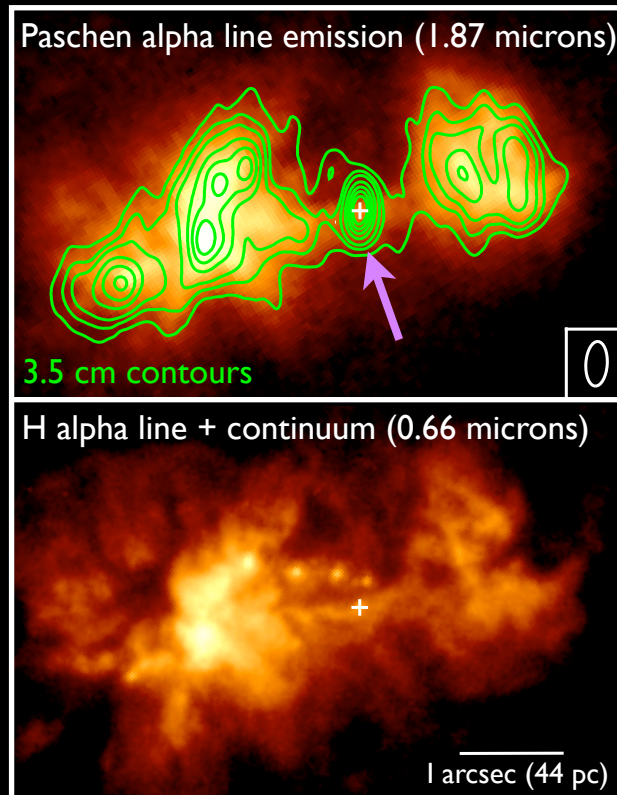
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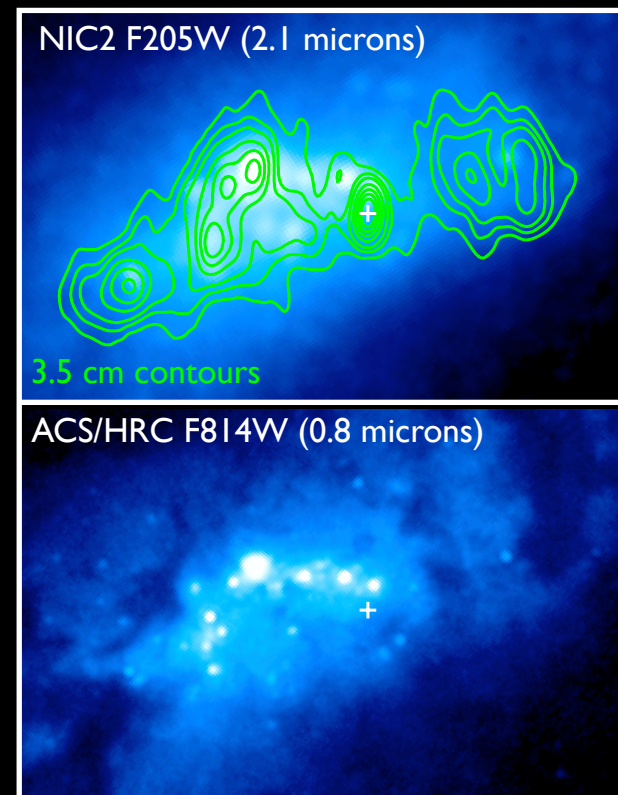
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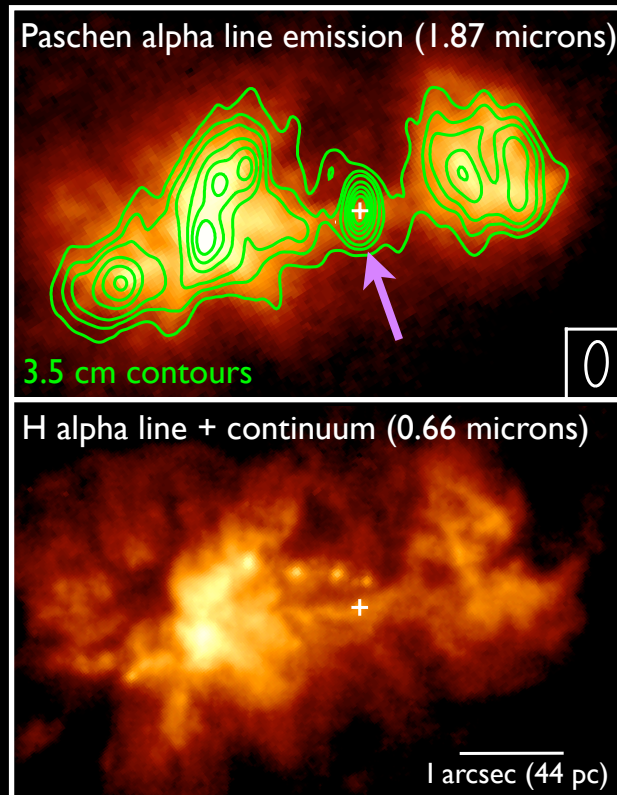
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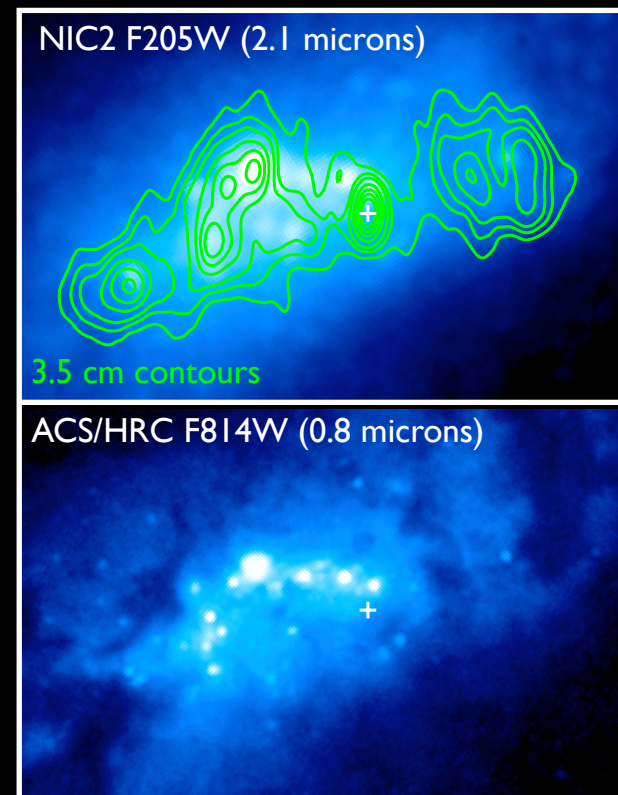
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→ **Active Galactic Nucleus**

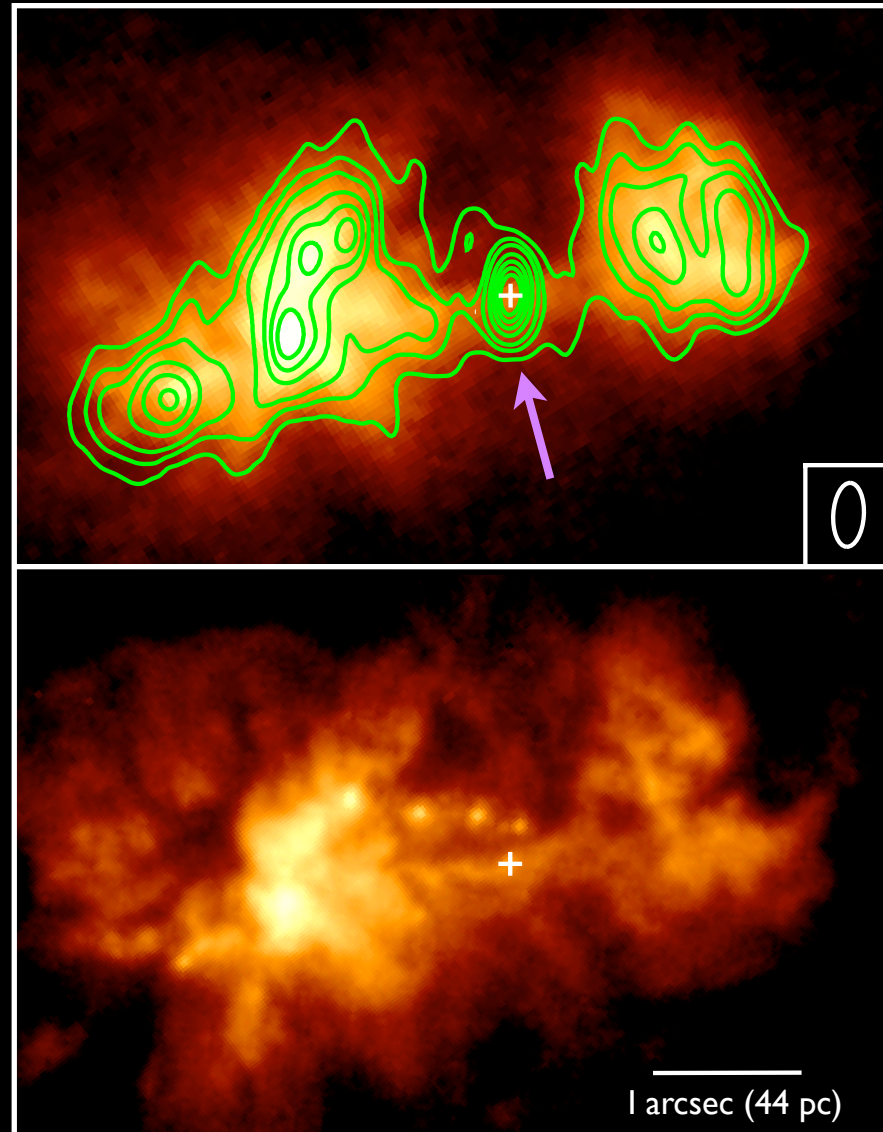


# Ruling out alternative explanations

- supernova remnants
- more recently created supernovae
- stellar mass black hole X-ray binaries
- some combination of the above

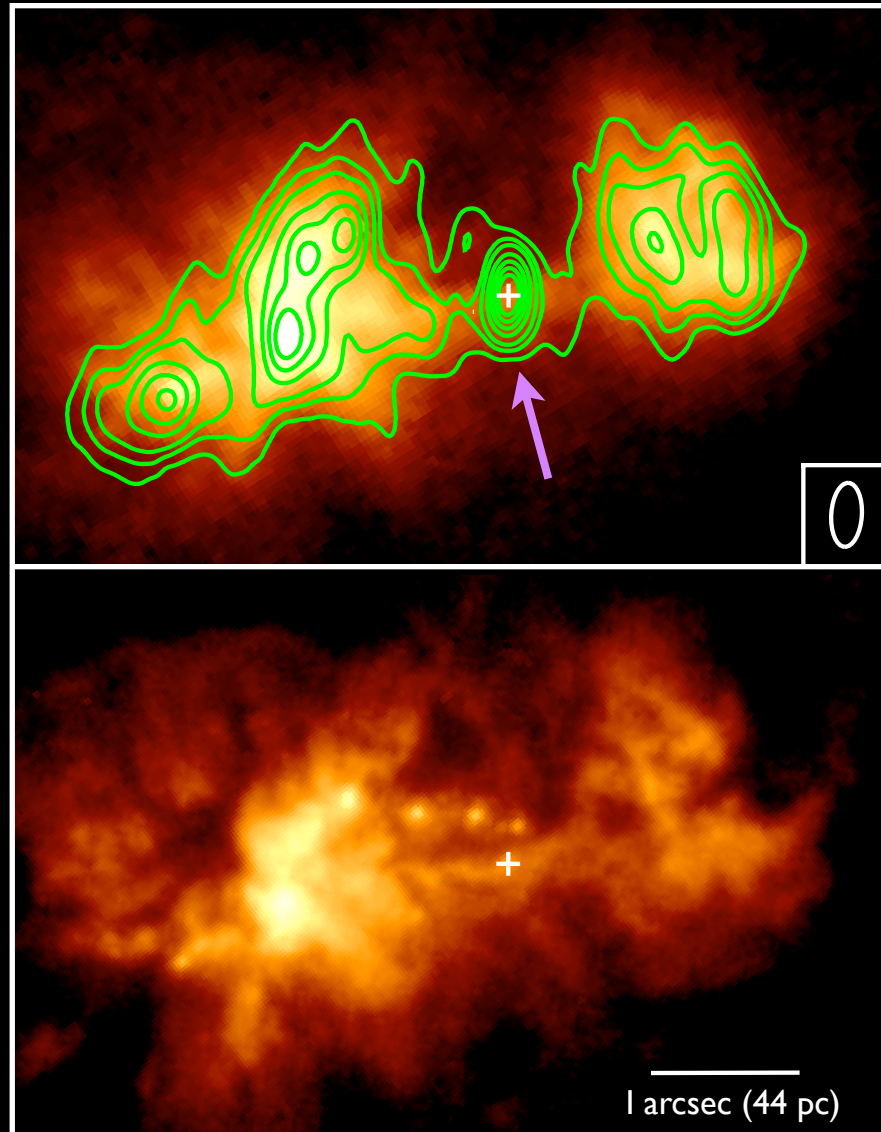
See Supplementary Information in Reines et al. 2011  
at [www.nature.com/nature](http://www.nature.com/nature) or on astro-ph

# An actively accreting massive black hole



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*How massive?*

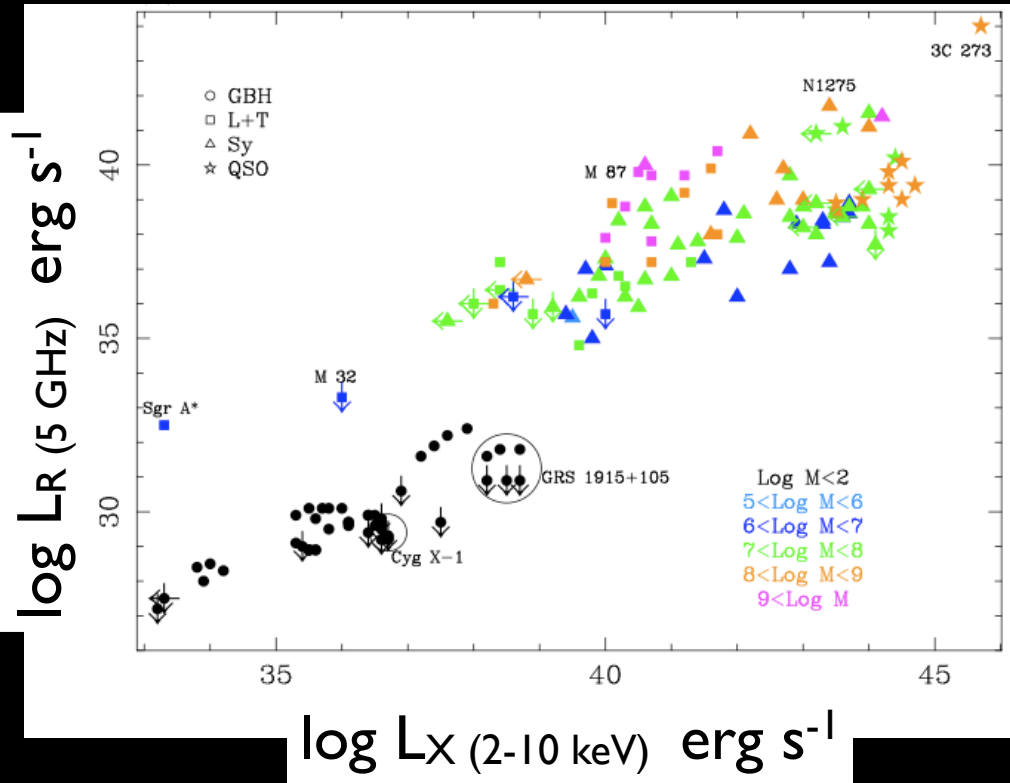




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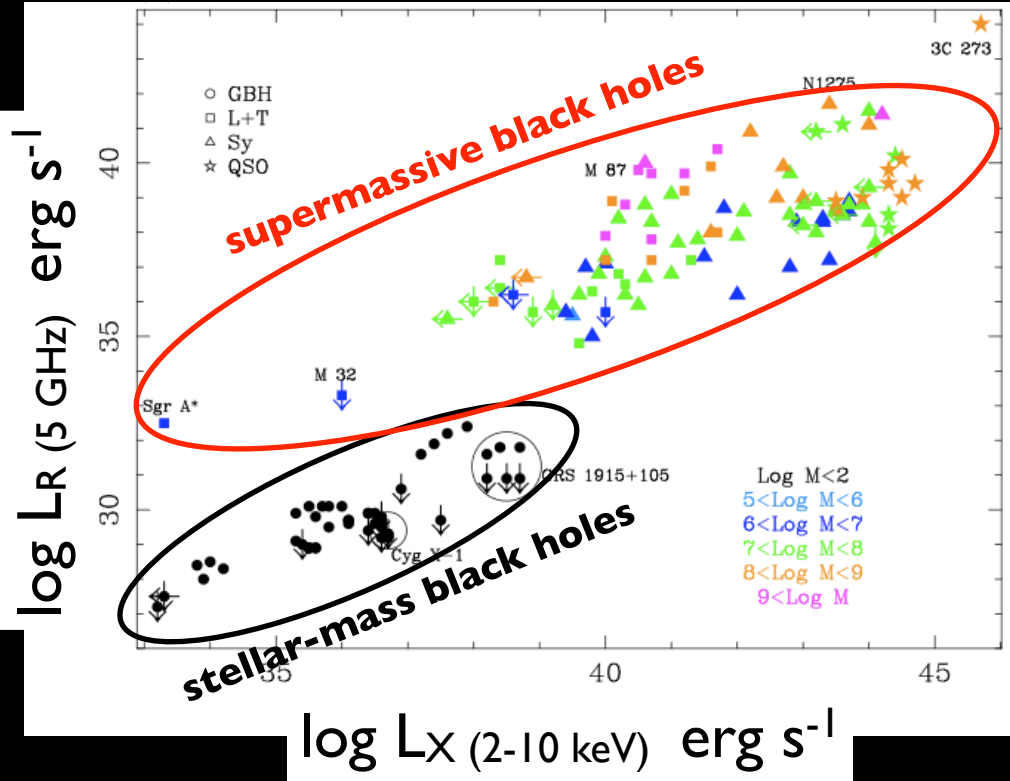
Merloni et al. 2003



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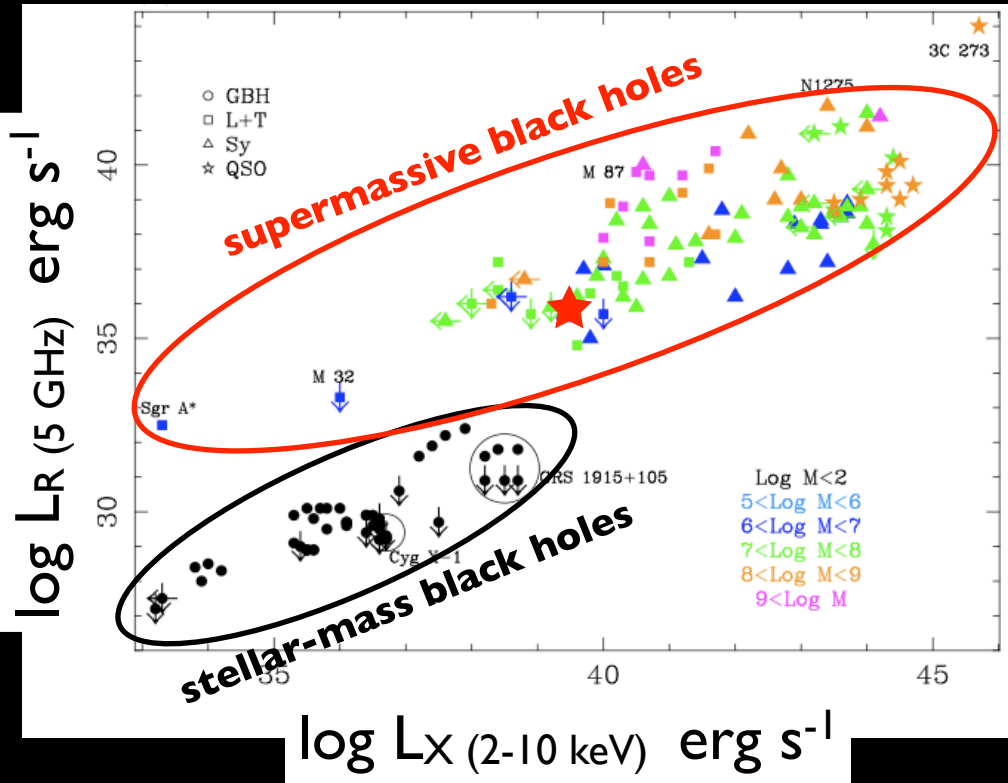
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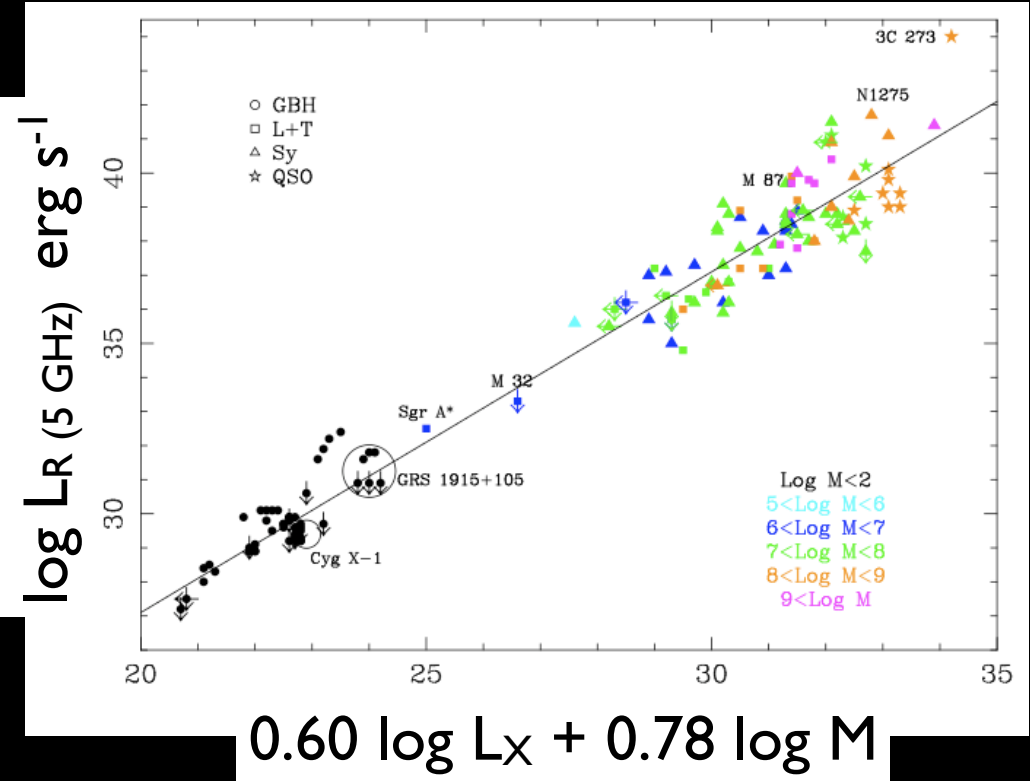
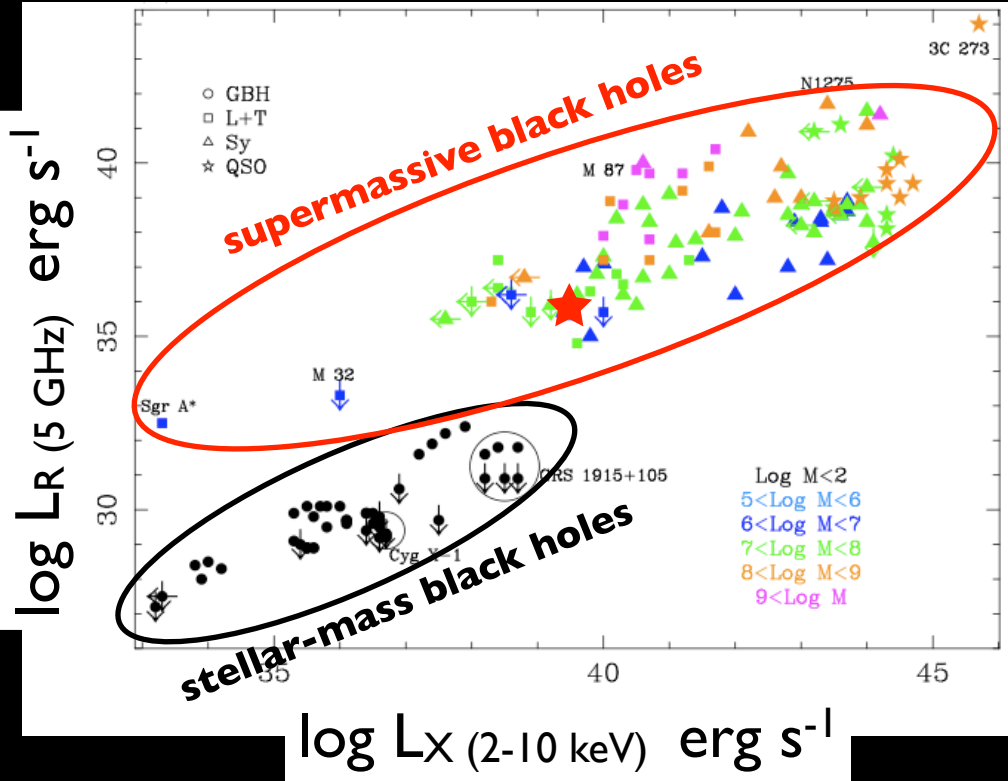




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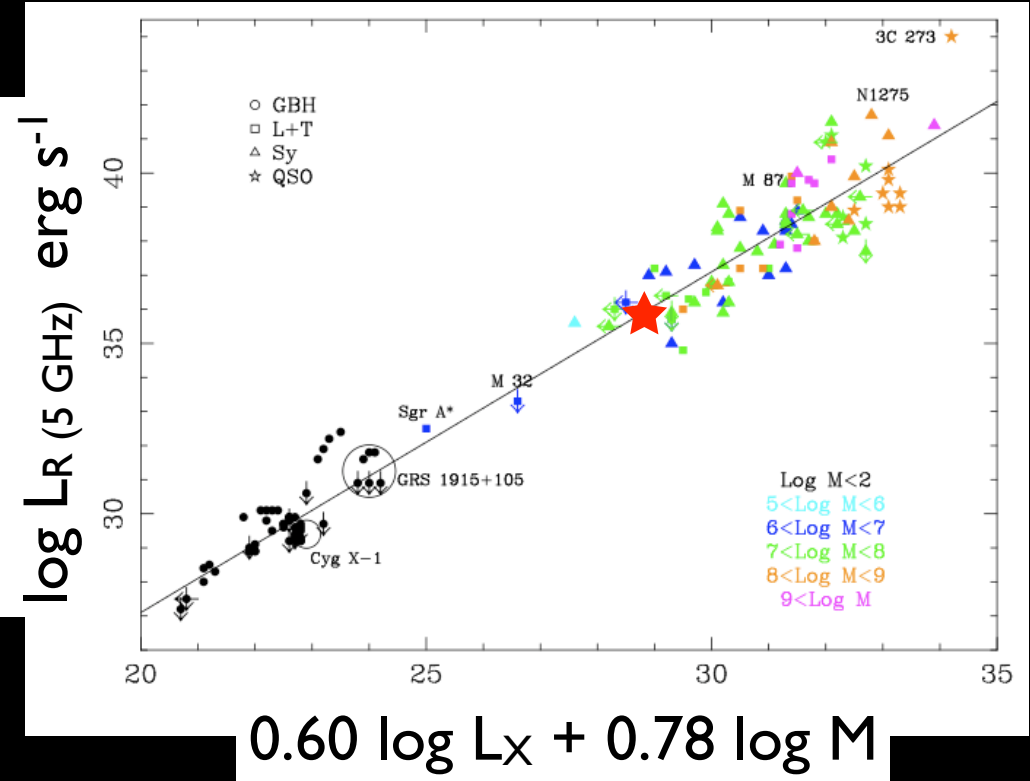
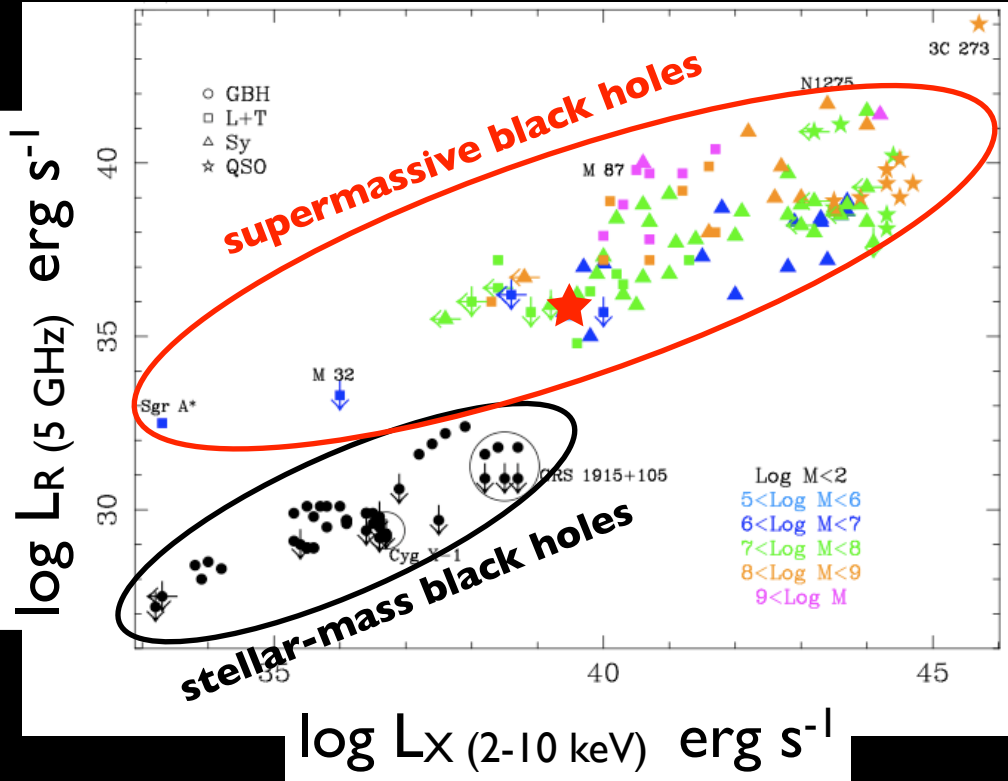
“fundamental plane of black hole activity”

$$\log L_R = 0.60 \log L_X + 0.78 \log M + 7.33$$

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*How massive?*

Merloni et al. 2003



“fundamental plane of black hole activity”

$$\log L_R = 0.60 \log L_X + 0.78 \log M + 7.33$$

black hole in Henize 2-10  $\longrightarrow \log (M_{\text{BH}}/M_{\text{sun}}) = 6.3 \pm 1.1$

Supermassive black holes have typically been found in massive galaxies with bulges



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*But not always...*



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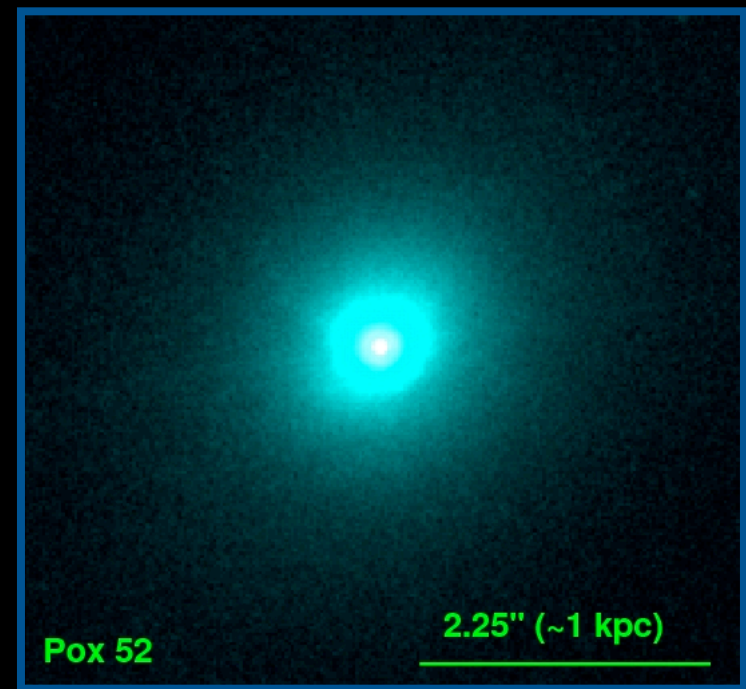
# The Low-Mass Regime: Putting Henize 2-10 in context

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Name	Type
NGC 4395	Sd
Pox 52	dE



Filippenko & Sargent (1989)  
Filippenko & Ho (2003)  
Peterson et al. (2005)

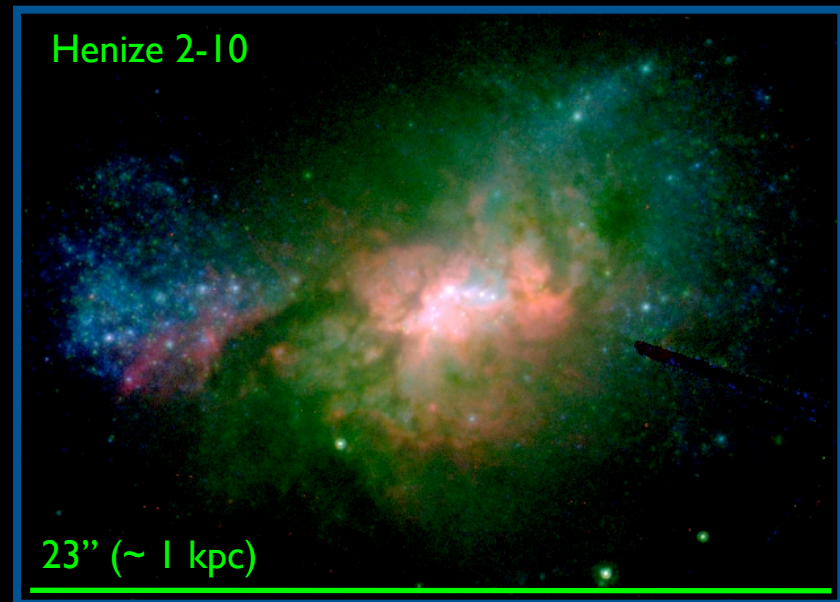


Kunth, Sargent & Bothun (1987)  
Barth et al. (2004)  
Thornton et al. (2008)

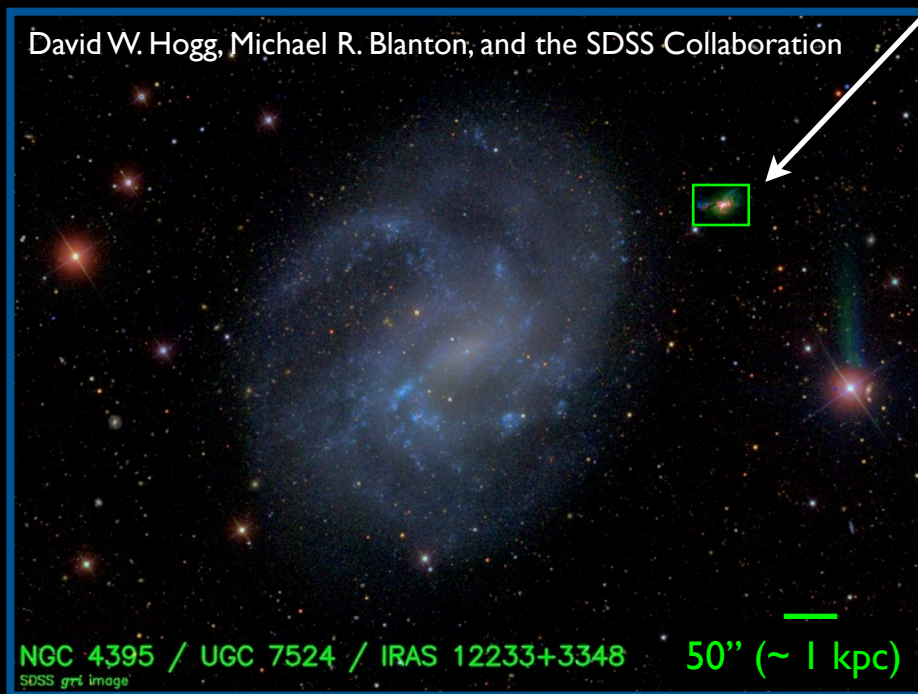


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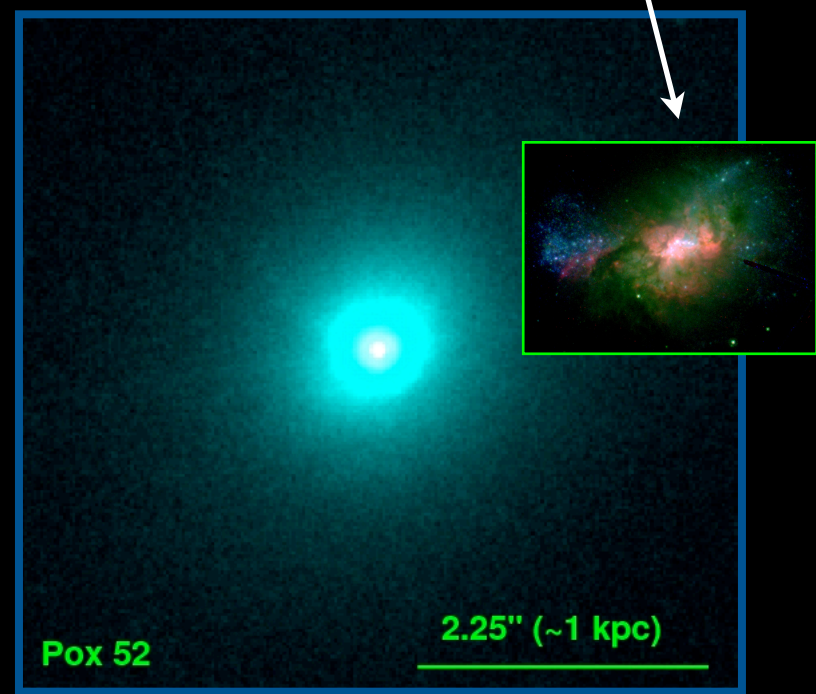
Name	Type
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Henize 2-10	Blue Compact Dwarf



Reines et al. (2011)



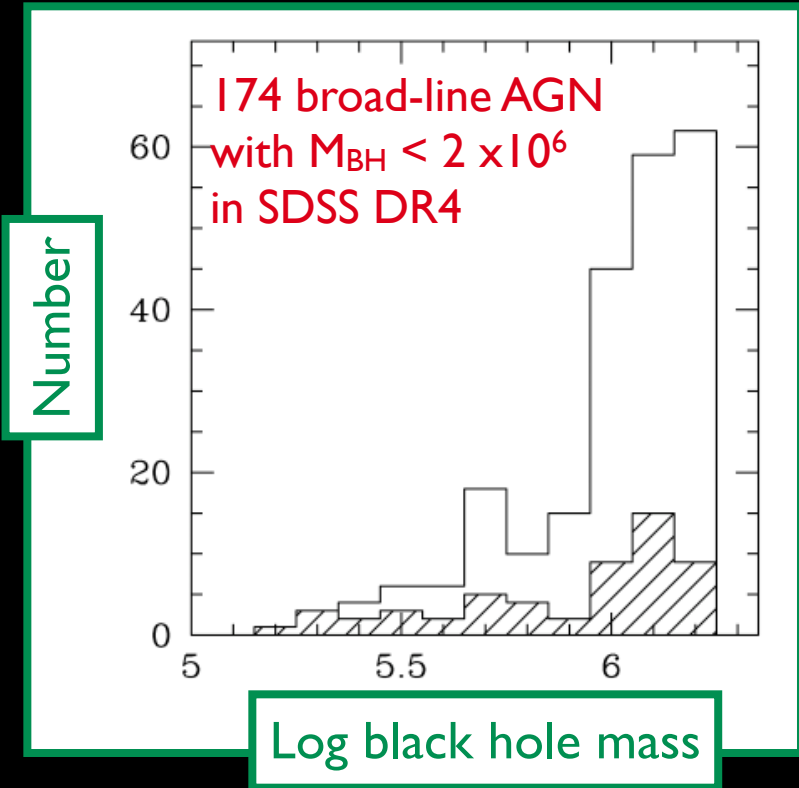
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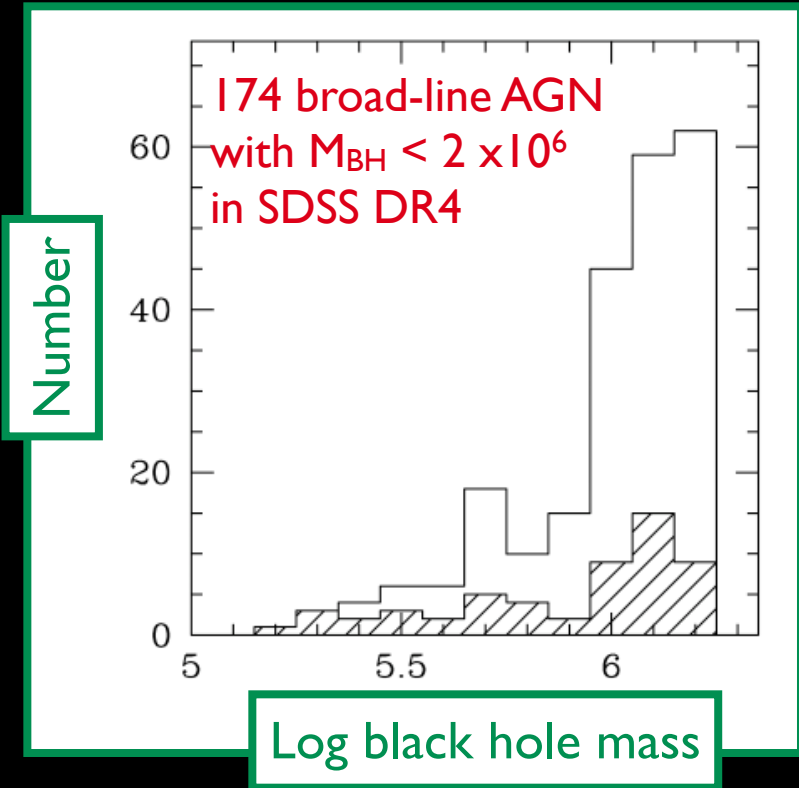
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Greene & Ho (2004, 2007)



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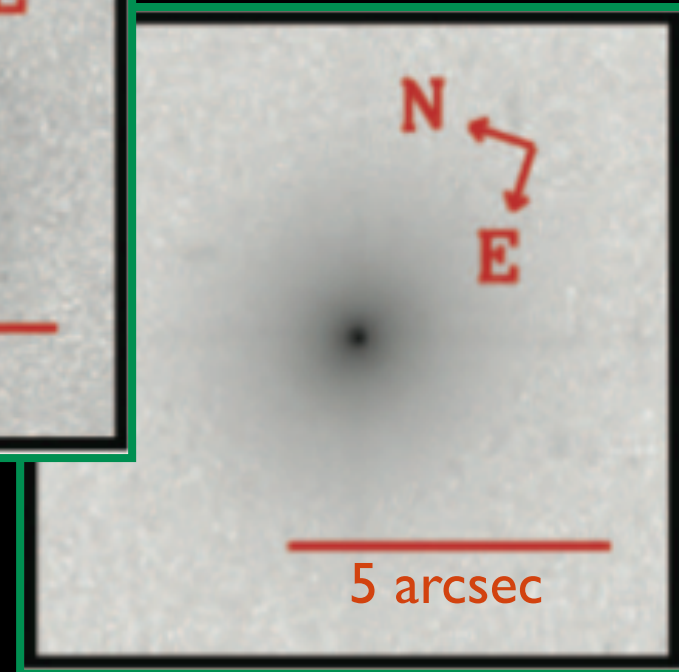
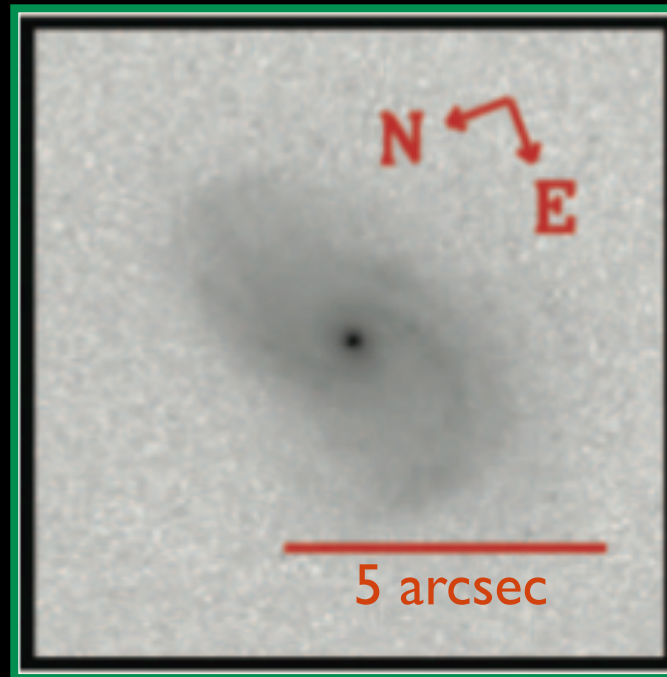
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~ 93% extended disks  
(with pseudobulges)

## Host Galaxies

- Low-luminosity galaxies,  $\sim 1$  mag below  $L^*$
- Well-defined optical nuclei

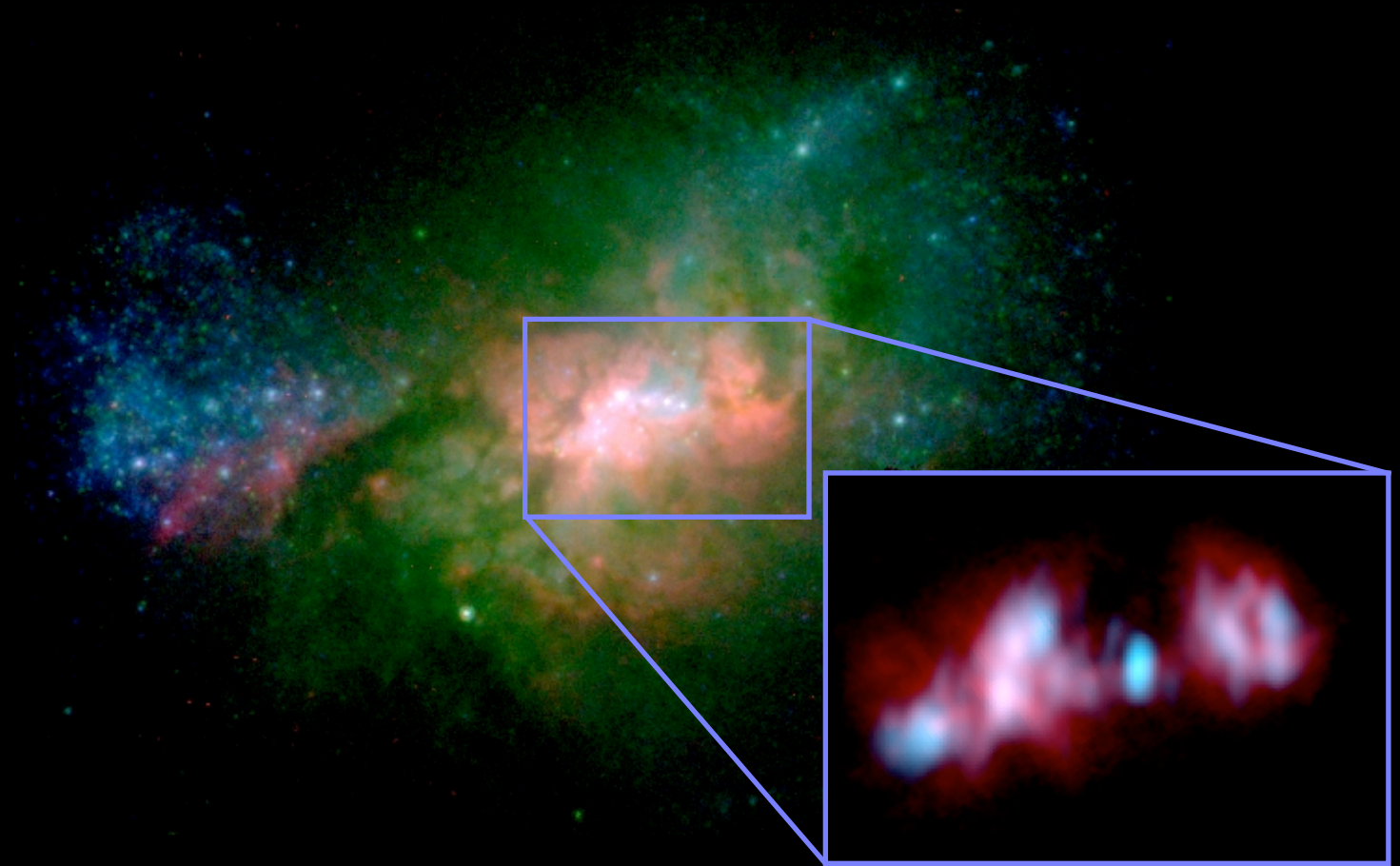


~ 7% spheroidals



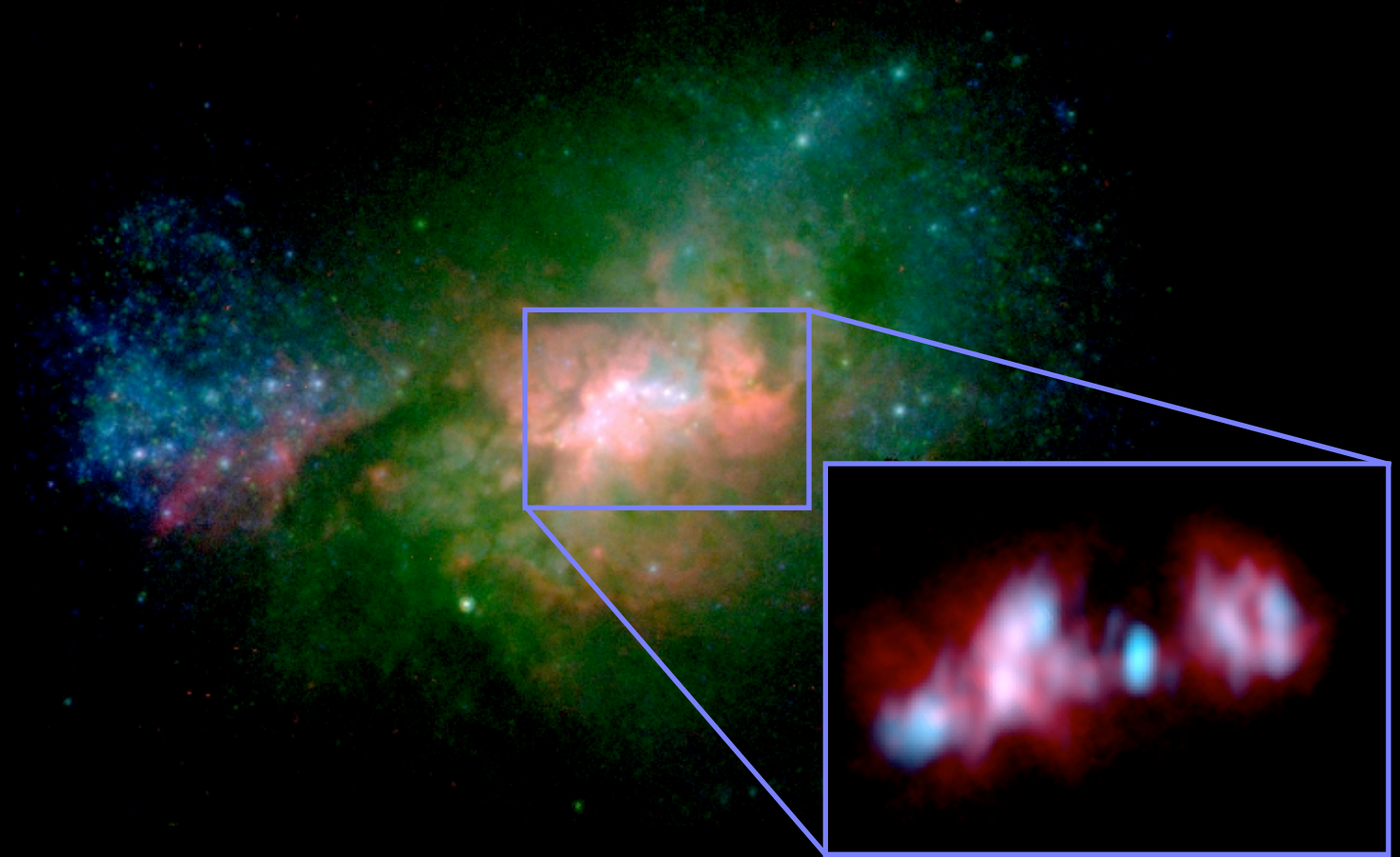
## Henize 2-10 is *different*

- Dwarf starburst galaxy with newly formed globular clusters
- Irregular morphology without a well-defined nucleus



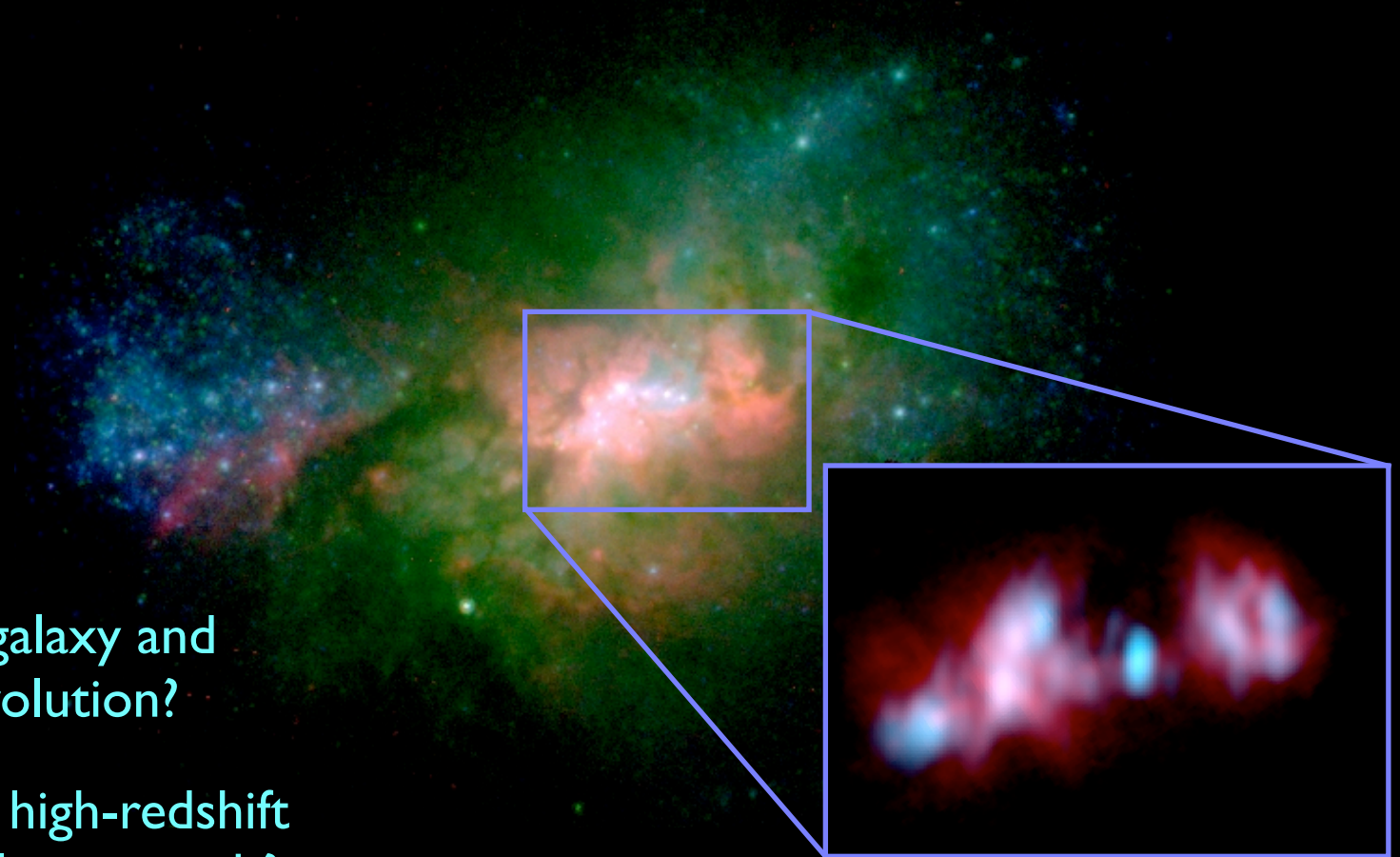
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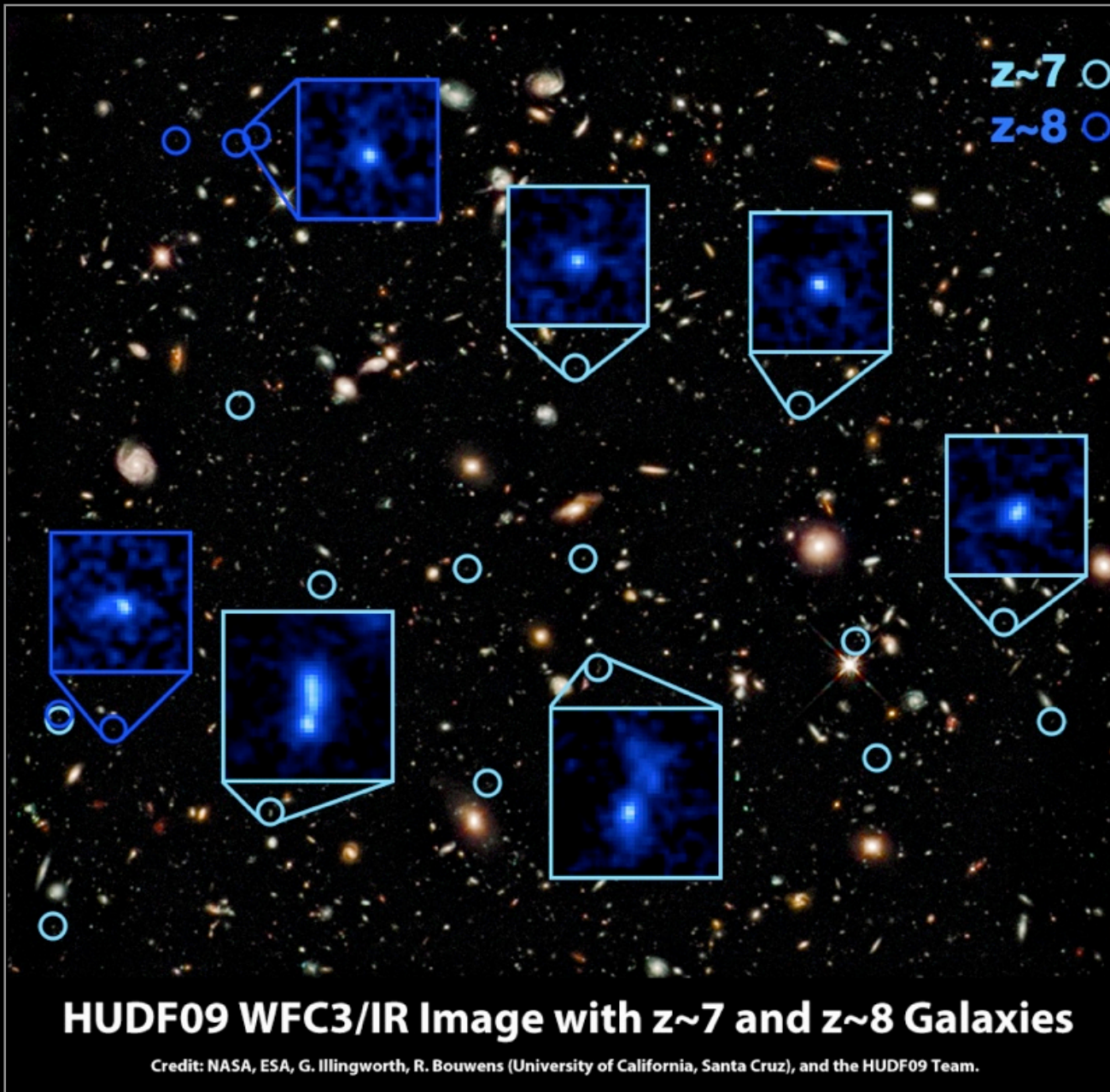


Early stage of galaxy and  
black hole evolution?

Local analogue to high-redshift  
black hole and galaxy growth?



# The First Star-Forming Galaxies

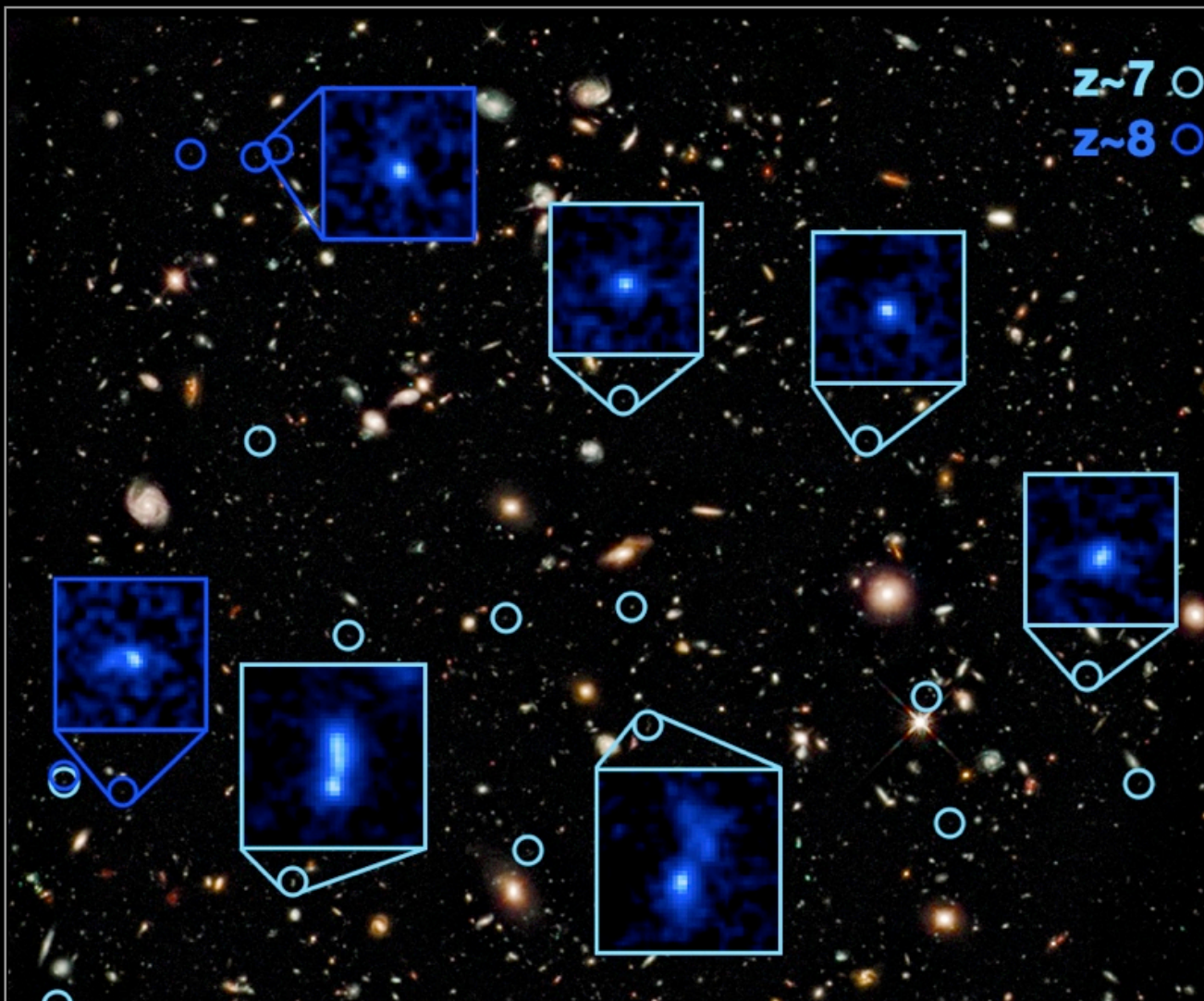


- blue, compact galaxies  
600-800 Myr after the Big Bang (Bouwens et al. 2010)
- intrinsic sizes  $\lesssim 1$  kpc  
(Oesch et al. 2010)
- masses  $\sim 10^9$ - $10^{10} M_{\text{sun}}$   
(Labbé et al. 2010)
- likely forming globular clusters
- likely host massive black holes (Treister et al. 2011)





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- likely forming globular clusters
- likely host massive black holes (Treister et al. 2011)

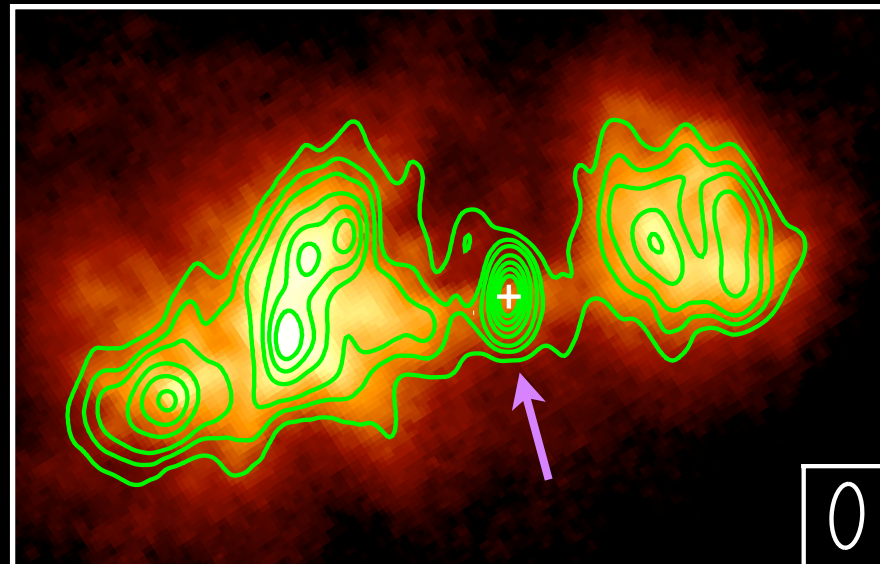


Henize 2-10 is our best available local analog of high-redshift black hole and galaxy growth

HUD

# Main take-away points about Henize 2-10

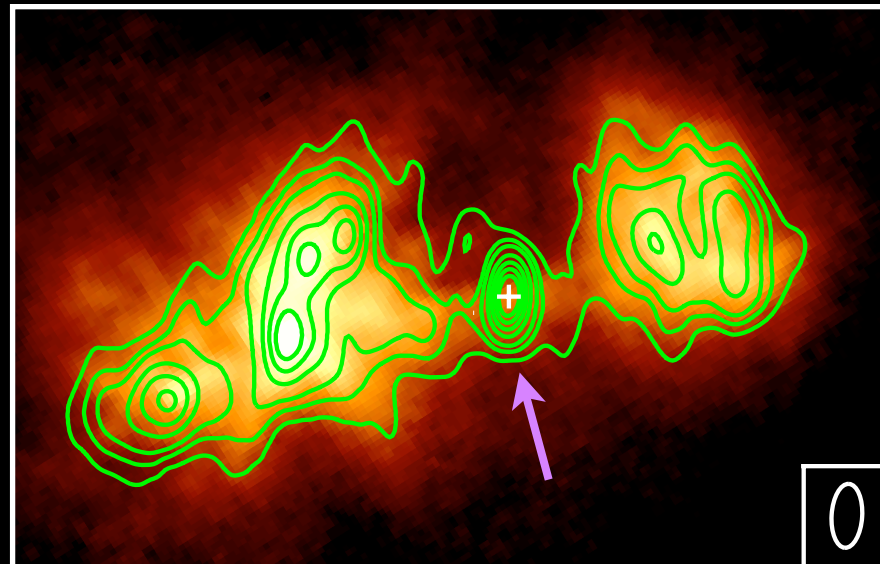
- First example of a massive black hole in a local dwarf starburst galaxy
- Nearby galaxy much like those in the earlier universe
- No discernible bulge - black hole growth can precede the build-up of galaxy spheroids





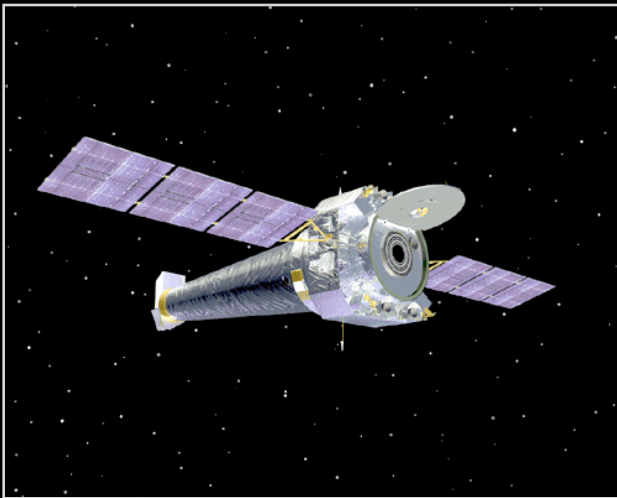
# Main take-away points about Henize 2-10

- First example of a massive black hole in a local dwarf starburst galaxy
- Nearby galaxy much like those in the earlier universe
- No discernible bulge - black hole growth can precede the build-up of galaxy spheroids
- Opens up a new class of host galaxies in which to search for low-mass black holes!



# The Future: Searching for big black holes in little galaxies

Sensitive, high-resolution radio and X-ray observations of star-forming dwarf galaxies



Accepted Cycle 13 *Chandra* + EVLA proposal

P.I. Reines (w/ Sivakoff, Condon)

# The Future: Searching for big black holes in little galaxies

*HST* and ALMA follow-up to study host galaxies





# Follow-up observations of Henize 2-10

## Accepted Proposals

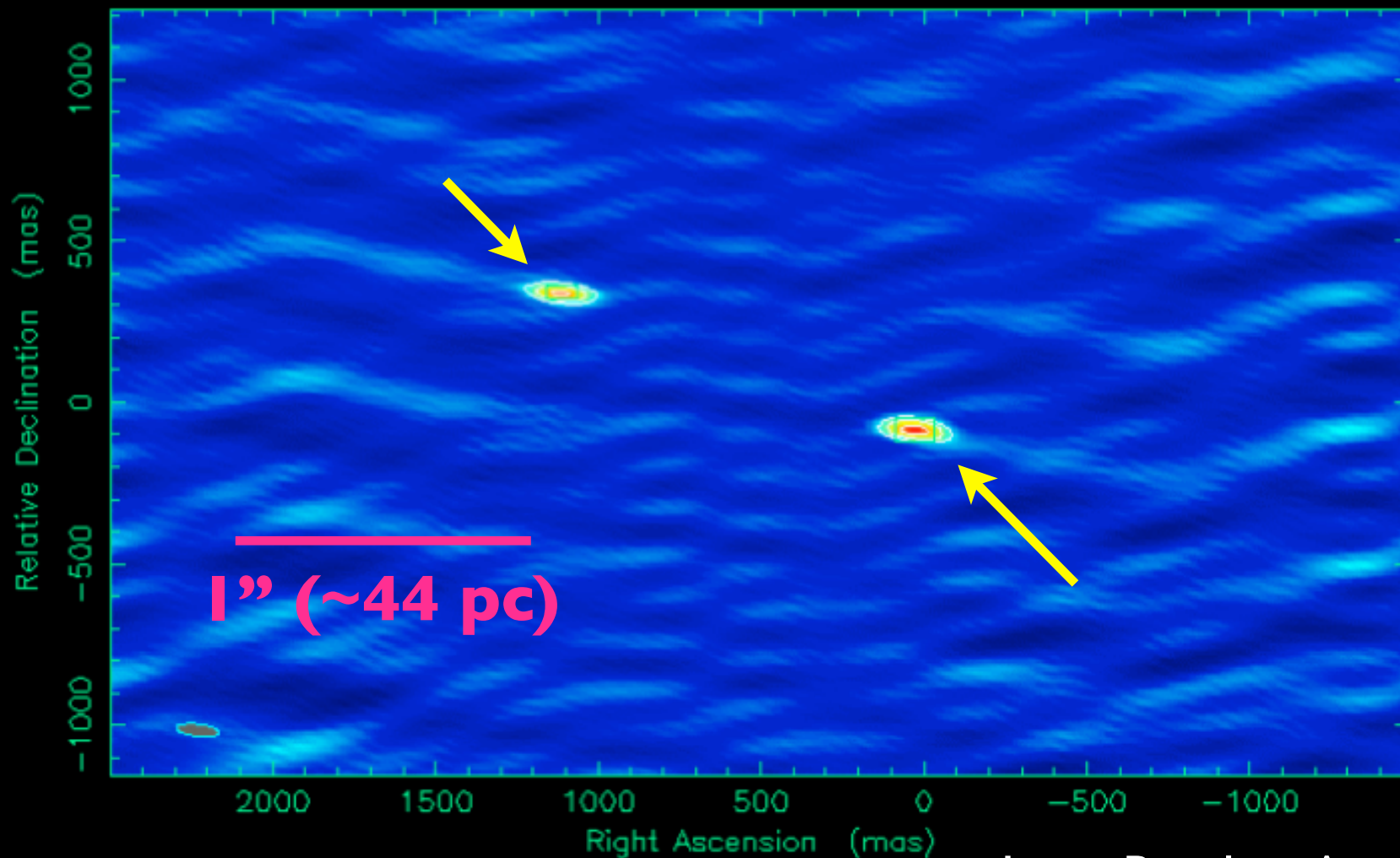
- *HST/STIS* - Kinematics and ionization conditions near AGN  
P.I. Reines (w/ Whittle, Johnson)
- *XMM-Newton* - X-ray follow-up  
P.I. Hickox (w/ Greene, Reines, Sivakoff, Johnson, Alexander)
- *VLBI with the Long Baseline Array* - High-resolution observations at 1.4 GHz  
P.I. Reines (w/ Deller, Johnson)

## Submitted Proposals

- *EVLA* - Water maser observations  
P.I. Reines (w/ Darling, Brogan, Johnson)
- *ALMA* - Dense molecular gas  
P.I. Johnson (w/ Reines, Testi, Brogan, Vanzi, Wilner, Chen)

# New VLBI observations of Henize 2-10 at 1.4 GHz

Clean I map. Array: AHMP AMP  
HE2-10 at 1.400 GHz 2011 Jul 22

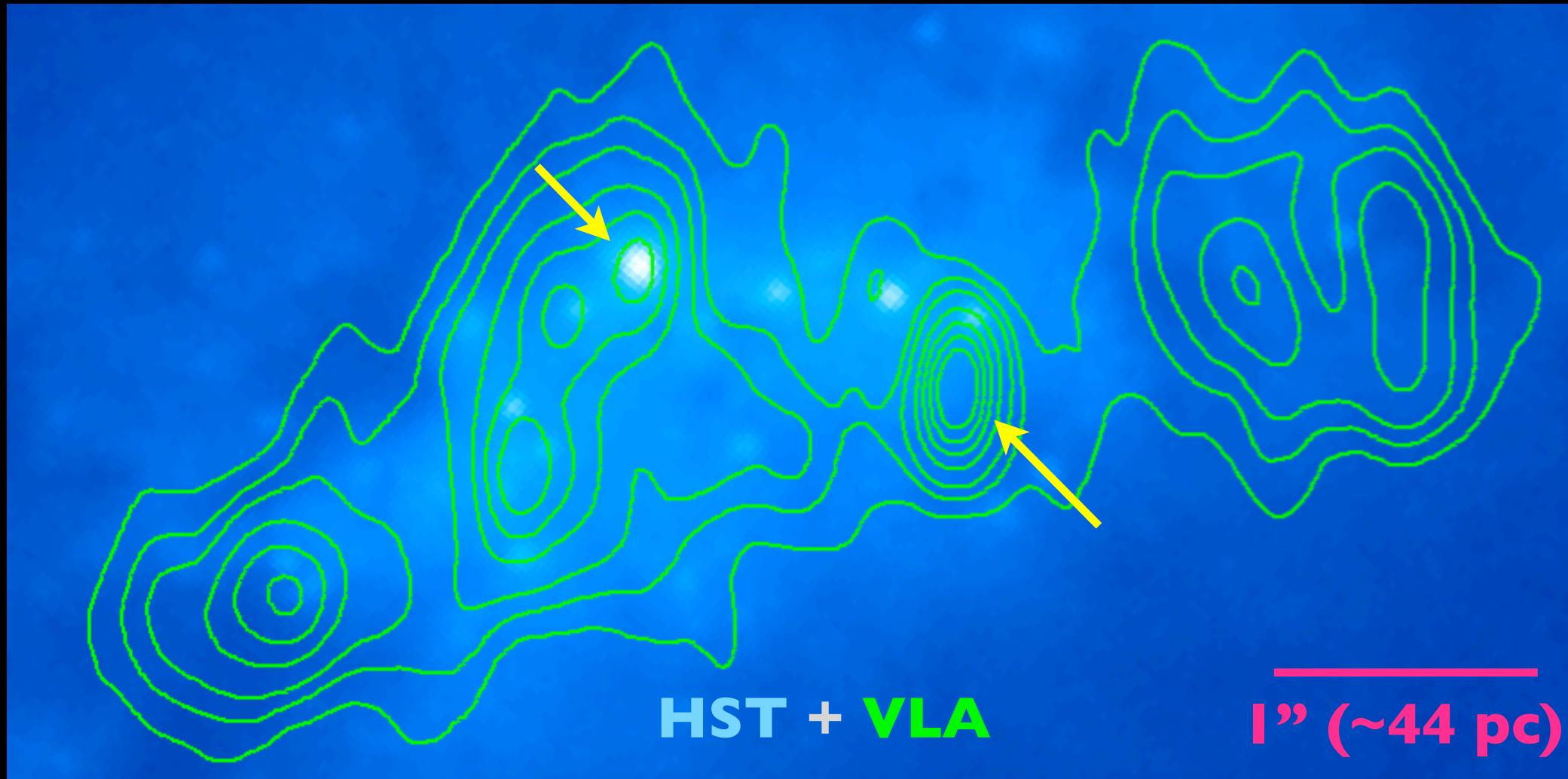


Map center: RA: 08 36 15.117, Dec: -26 24 34.070 (2000.0)  
Map peak: 0.000616 Jy/beam  
Contours %: 20 40 80  
Beam FWHM: 131 x 38 (mas) at 83.1°

Long Baseline Array (LBA)  
w/ Adam Deller



# New VLBI observations of Henize 2-10 at 1.4 GHz





# New VLBI observations of Henize 2-10 at 1.4 GHz

SNR in brightest super star cluster

AGN

HST + VLA + LBA

1" (~44 pc)

