

Black hole mergers via clusters

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Motivation

- Three ingredients:

- Mass segregation --> many BH mergers

- [Portegeis Zwart & McMillan 2000, ApJL 528 17]

- [O'Leary et al, astro-ph/0508224]

- Many stars form in massive clusters

- [cluster IMF $\sim M^{-2}$; **all** stars form in clusters]

- Star formation history

- (...doesn't matter much)

new

Simple explanation

See poster!

Simple explanation: Assumptions

- Thermal BH-BH binaries
- $M_{\text{BH}} = 10 M_{\text{O}}$

Simple explanation: Ingredients

- Merger rate density

- Per mass

- ...in massive cluster

$$R(t) = 2 \times 10^{-5} M_{\odot}^{-1} / t$$

- Overall, at ‘t’

$$\mathcal{R}(t) = \int_{-T_{uni}}^t R(t - \tau) \frac{d\rho_{cl}}{dt} d\tau$$

- Star formation history

- + fraction massive enough ($\sim 1/3$)

- + fraction surviving “infant mortality”

$$\left. \begin{array}{l} + \text{fraction massive enough } (\sim 1/3) \\ + \text{fraction surviving "infant mortality"} \end{array} \right\} \frac{d\rho_{cl}}{dt} = f_{cl} \frac{d\rho_{SFR}}{dt}$$

Simple explanation: Results

- Early only

- [Order of magnitude]
- Starburst at 13 Gyr

$$\frac{d\rho_{cl}}{dt} = f_{cl}\rho_{stars}\delta(t - 13\text{Gyr})$$

Density in stars:

10% (stars/baryon) * 4%(baryons/total)
* closure density

$$\sim 2 \times 10^8 \text{ M}_\odot/\text{Mpc}^3$$

Merger rate:

$$\sim 3 \times 10^{-7} (3f_{cl})\text{Mpc}^{-3}/\text{yr}$$

LIGO-I detections?

...range ~ 100 Mpc

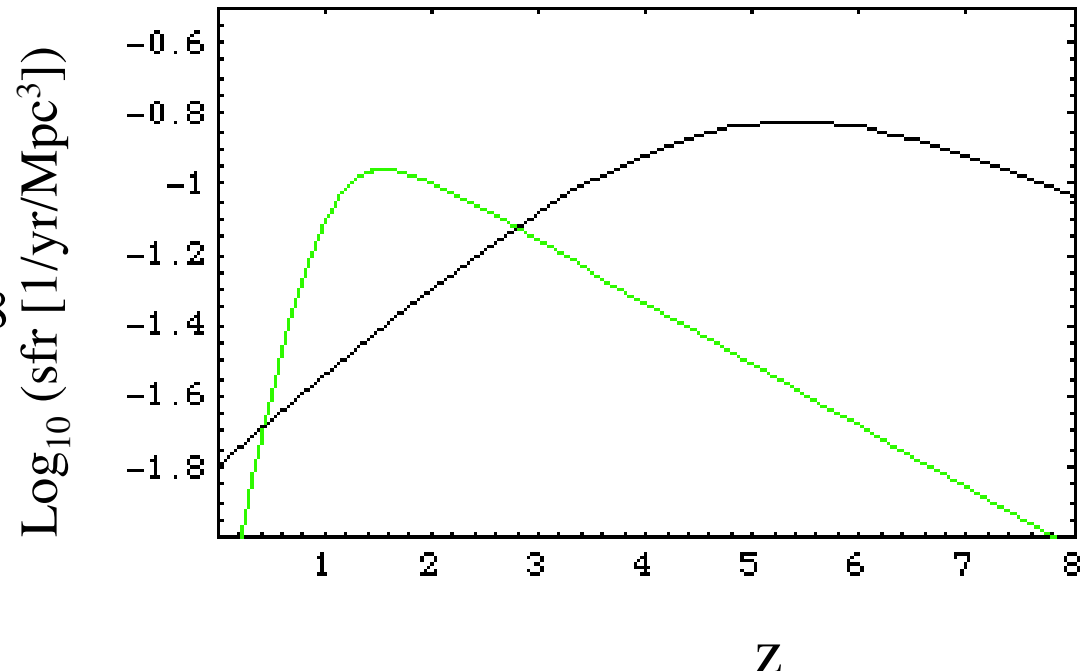
$$\sim 1.6 (3f_{cl})/\text{yr}$$

Simple explanation: Results

- Realistic SFRs:

Models

- Experimental
Madau, astro-ph/9907268
- Theory
Springle & Hernquist
MNRAS 339 312



Results [preliminary]

LIGO rate $\sim 5 \quad (3 f_{cl})/\text{yr} \quad [\text{SH}]$

$\sim 1.8 \quad (3 f_{cl})/\text{yr} \quad [\text{Madau}]$

- Clusters which form **recently** dominate ($\sim 80\%$)
...recent SFR matters most (b/c of $1/t$)

Detailed results

- Ingredients:
 - Cluster/ejection modelling
 - ...not demanding 1/t a priori
 - realistic BH IMF
 - ... heavier BHs...see **farther**...higher rate
 - LIGO sensitivity vs frequency
 - ...some BH-BH mergers too massive to see
 - cosmology [LIGO-II]

Detailed results

- Results: (preliminary; network [H1+L1+H2] range)

	LIGO-I (yr)	LIGO-2 (yr)
S-H	1.4 ($3 f_{cl}$)	2.2×10^4 ($3 f_{cl}$)
Madau	0.81 ($3 f_{cl}$)	2.0×10^4 ($3 f_{cl}$)

via latest LIGO-I (now!) and LIGO-II sensitivity

MODEST significance

- as test of cluster dynamics
 - cluster formation, survival fraction, mass segregation
- as probe of **birth conditions** of clusters
 - ...eventually, via LIGO-II