

Extreme Solar Systems II, Jackson Lake, 12-16 September 2011

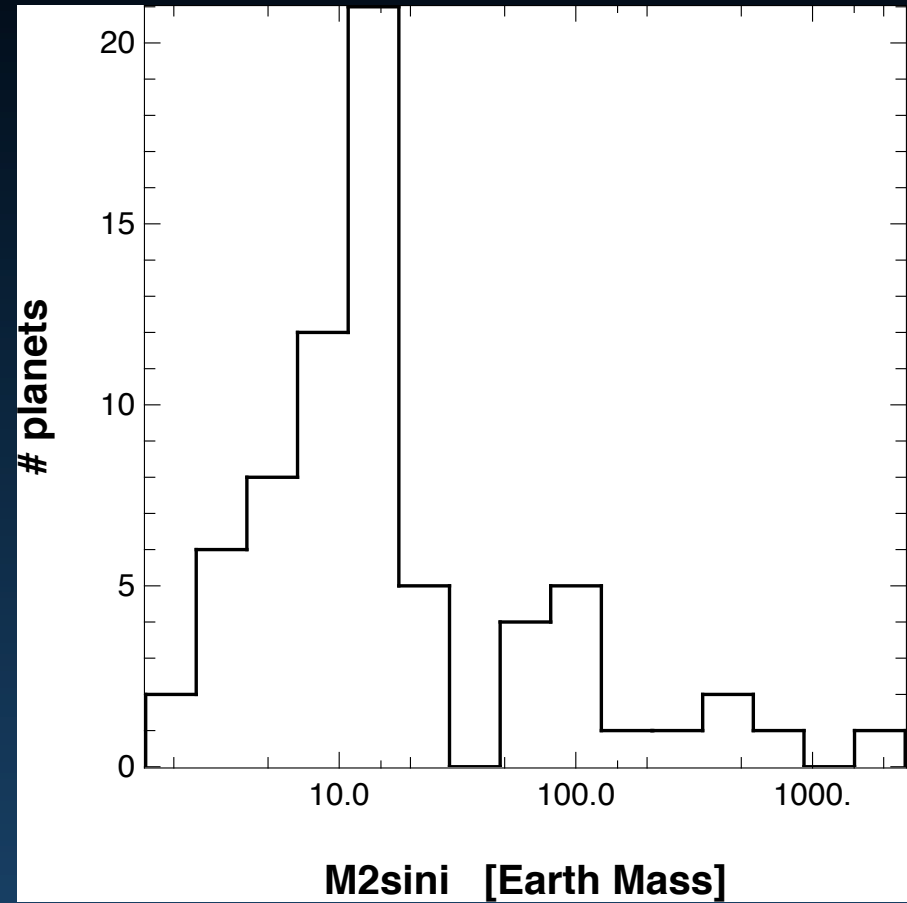
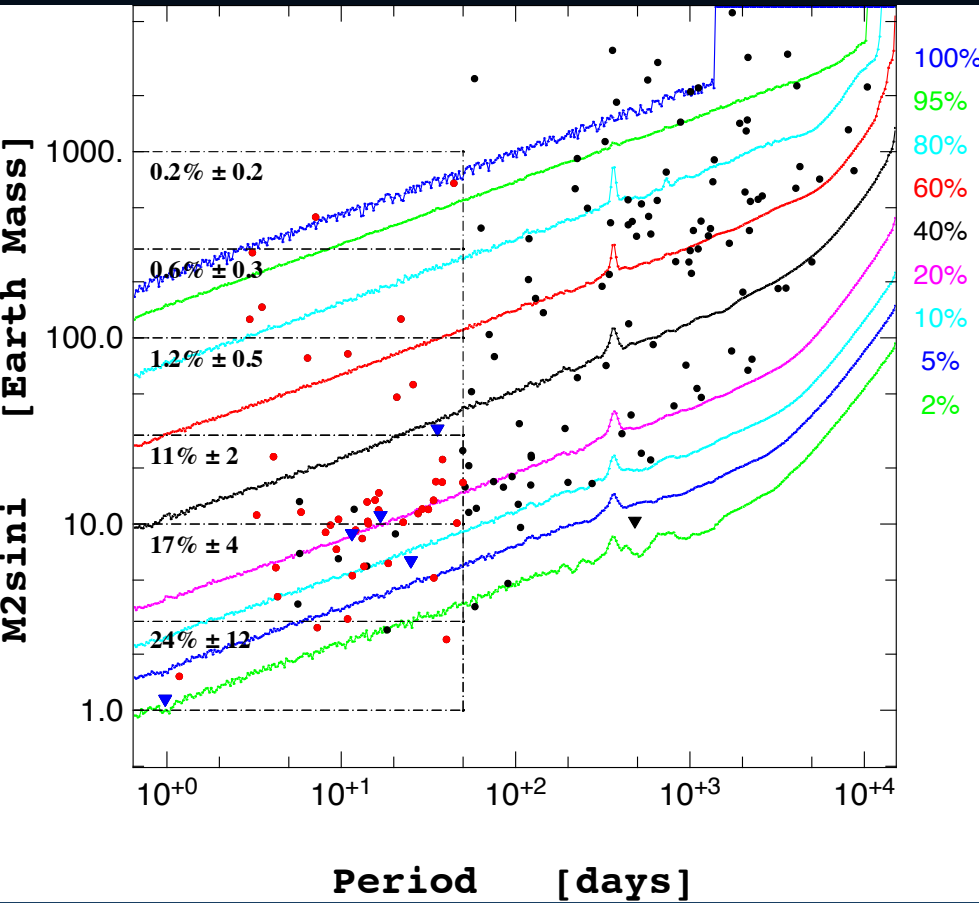
HARPS Long-Term Results: Effects of Stellar Magnetic Cycles, and Planet Detection Limits

Christophe Lovis
University of Geneva

in collaboration with:

X. Dumusque, N.C. Santos, F. Bouchy, M. Mayor, F. Pepe, D. Queloz, D. Ségransan, S. Udry

HARPS & CORALIE global results

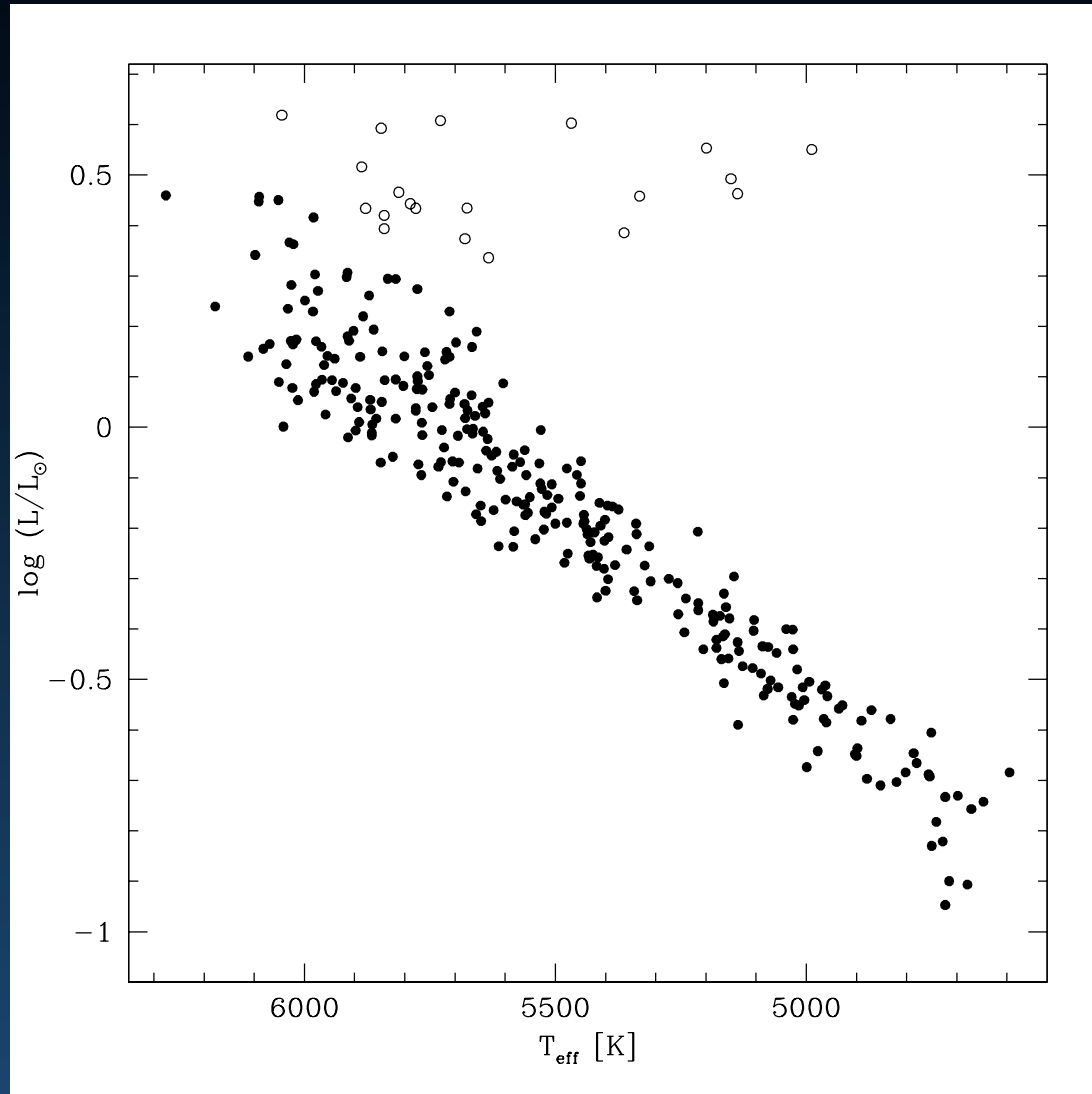


See Michel Mayor's talk!

Challenges to detect the low-mass planet population

- Instrument precision: reaching < 1 m/s is fundamental
- Stellar RV noise at different timescales:
 - p-mode oscillations, granulation (Dumusque et al. 2011, A&A 525, 140)
 - rotational modulations (Dumusque et al. 2011, A&A 527, 82)
 - magnetic cycles (Santos et al. 2010, A&A 511, 54, Lovis et al., arXiv:1107.5325, Dumusque et al., arXiv:1107.1748)

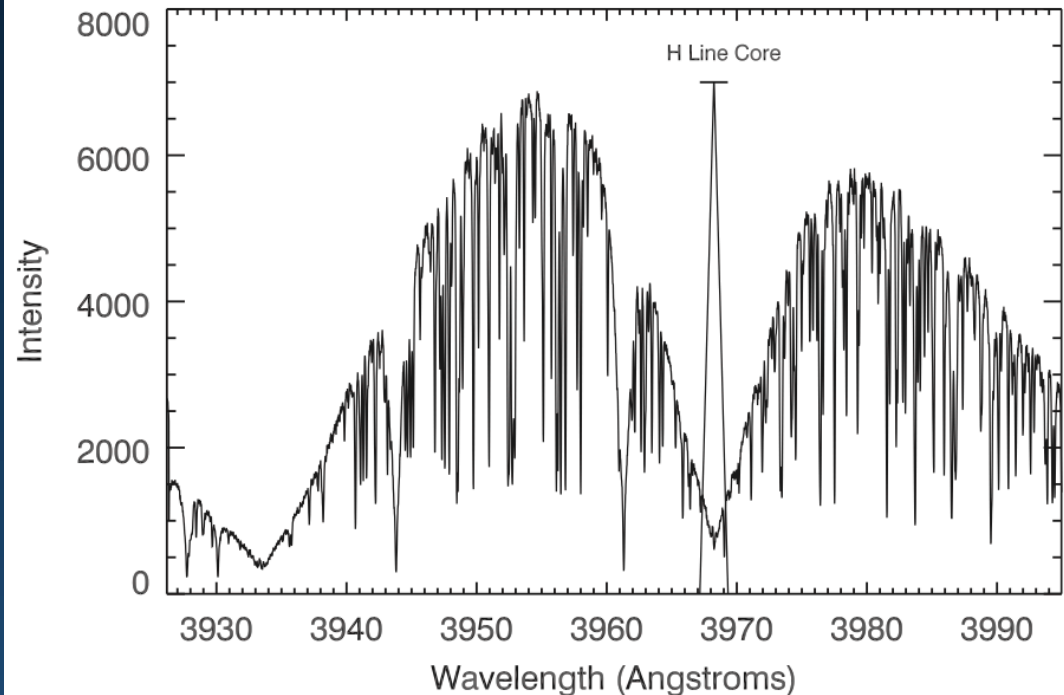
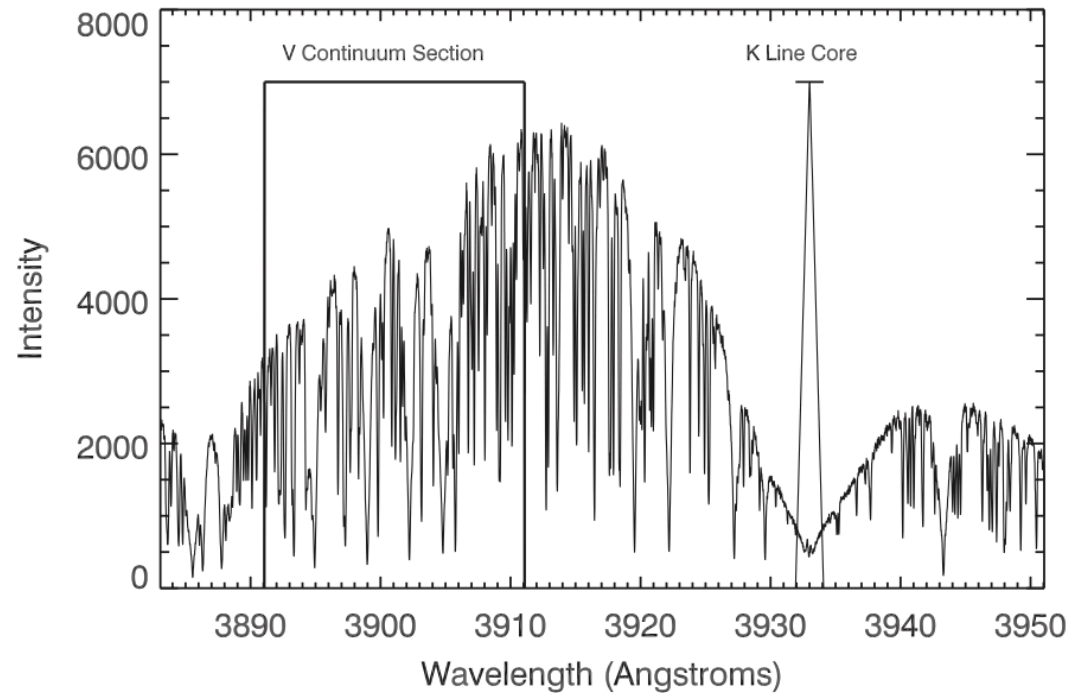
Stellar sample: 304 old FGK stars



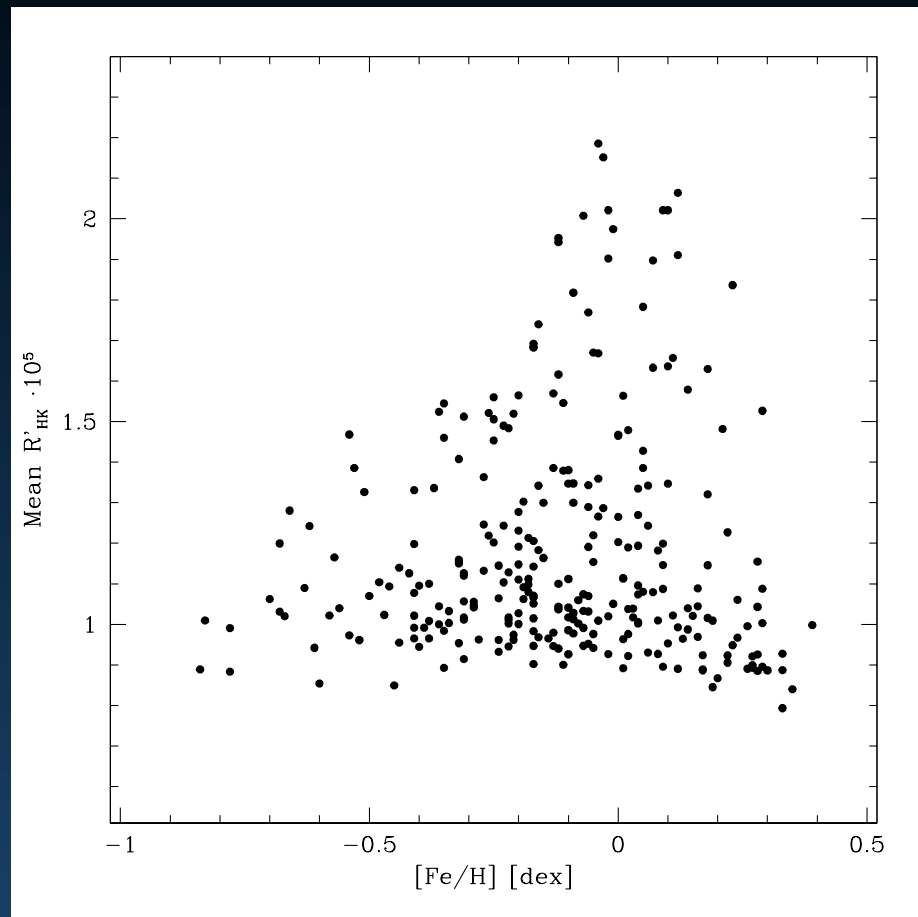
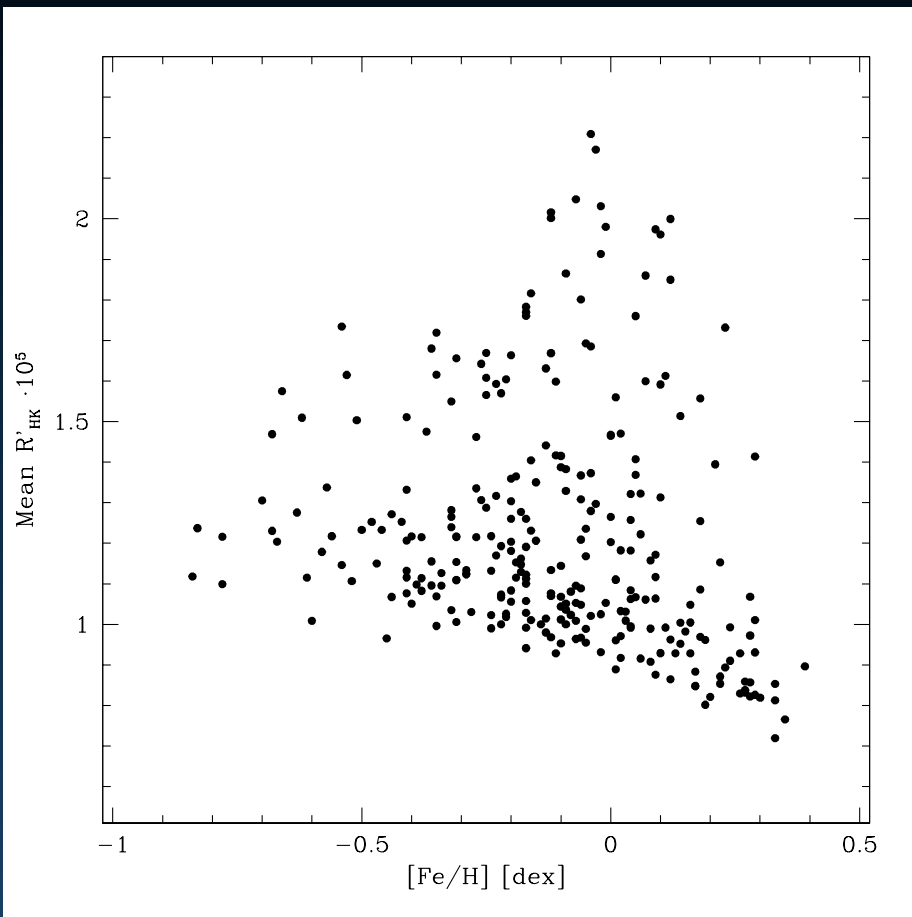
Ca II H&K lines
Mt Wilson S index
 R'_{HK} activity index

$$S = \alpha \frac{H + K}{R + V}$$

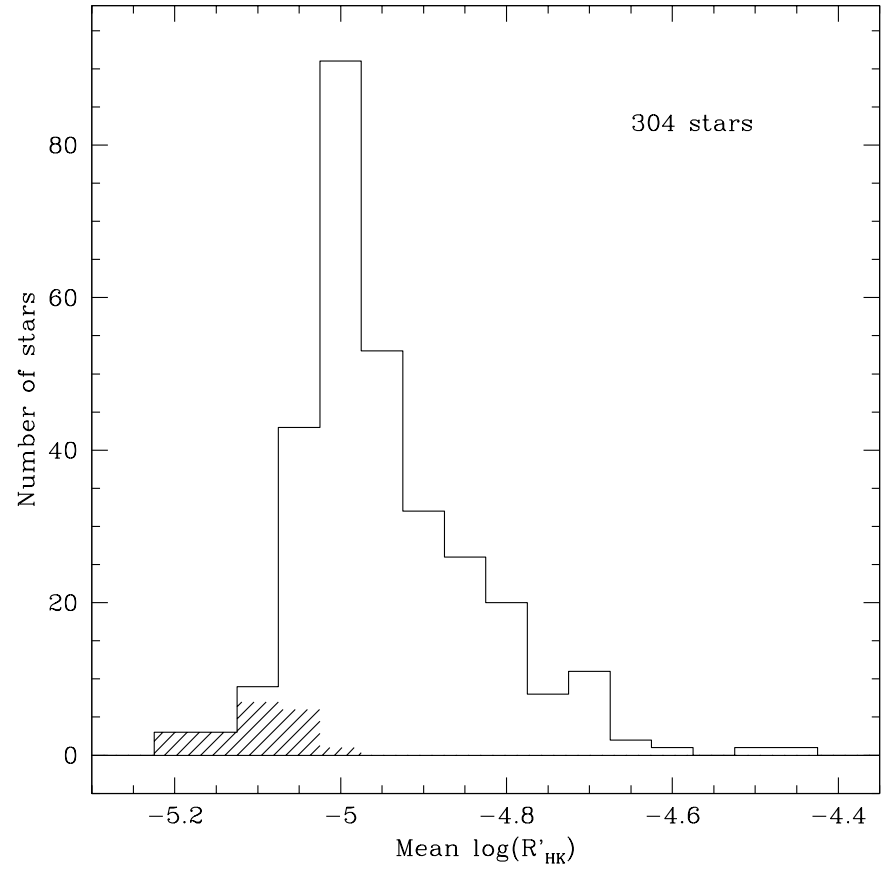
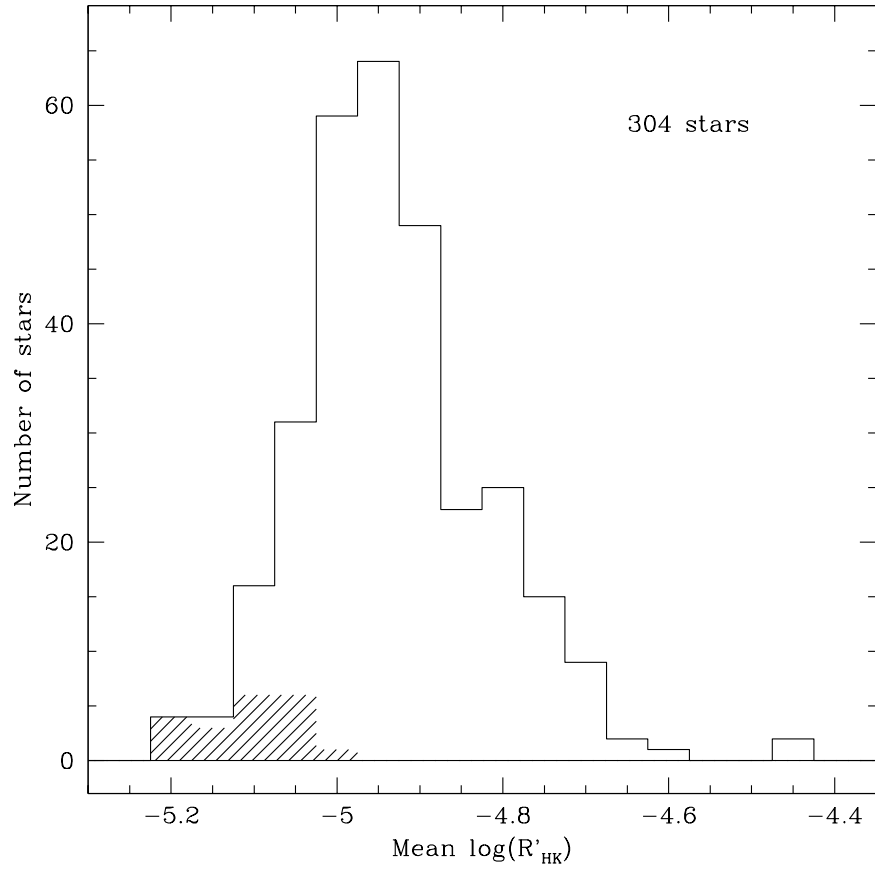
$$R'_{HK} = C_{cf}(B - V) \cdot S - R_{phot}(B - V)$$



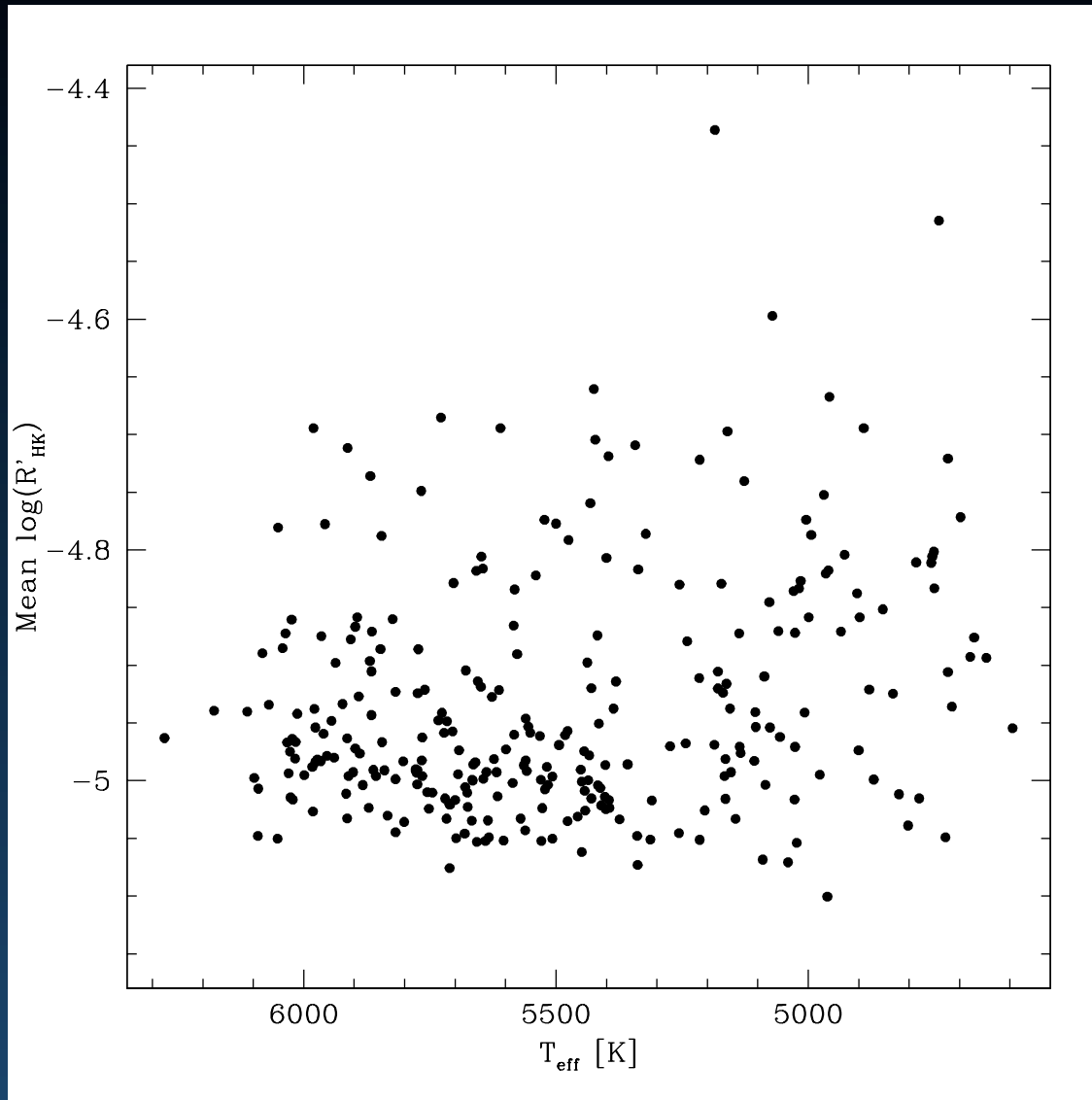
Effects of metallicity on R'_{HK}



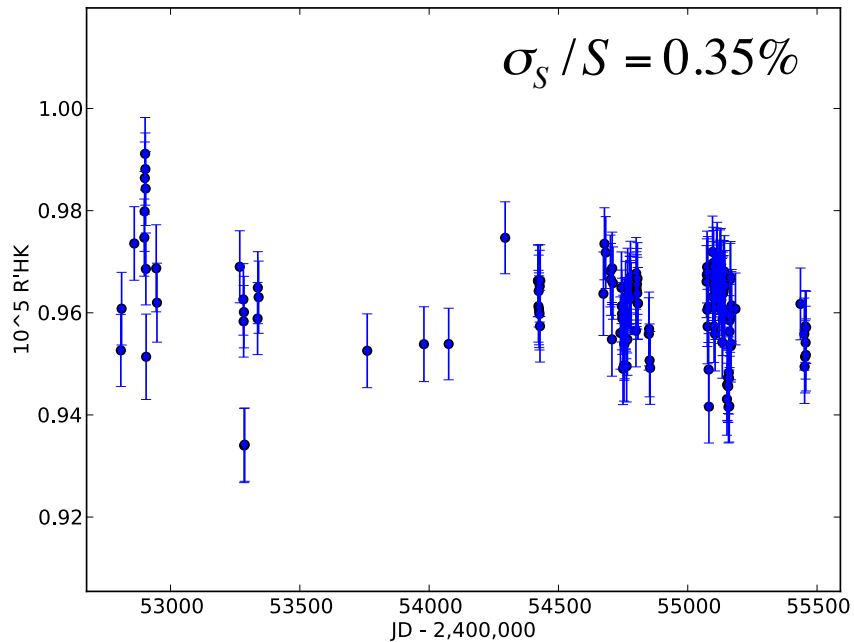
Effects of metallicity on R'_{HK}



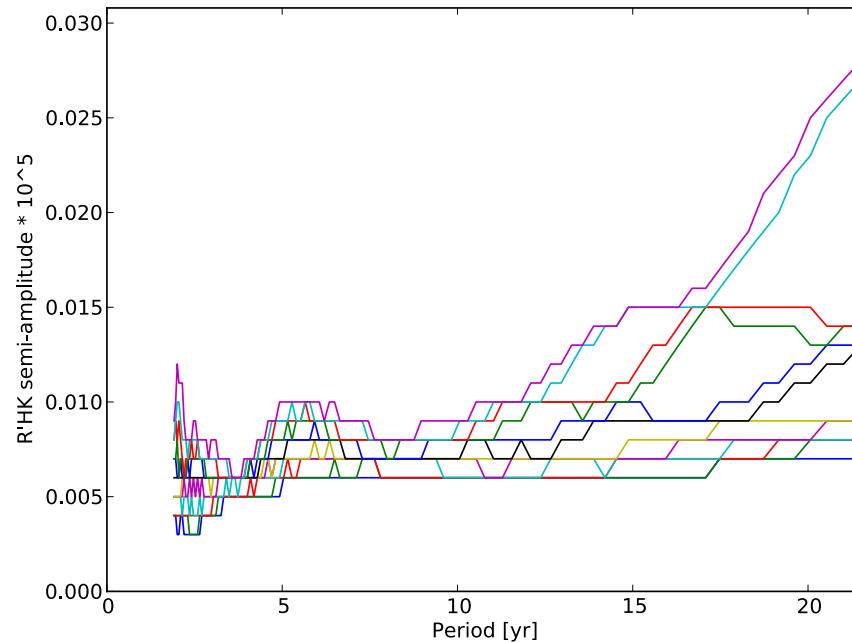
$\log(R'_{\text{HK}})$ vs. T_{eff}



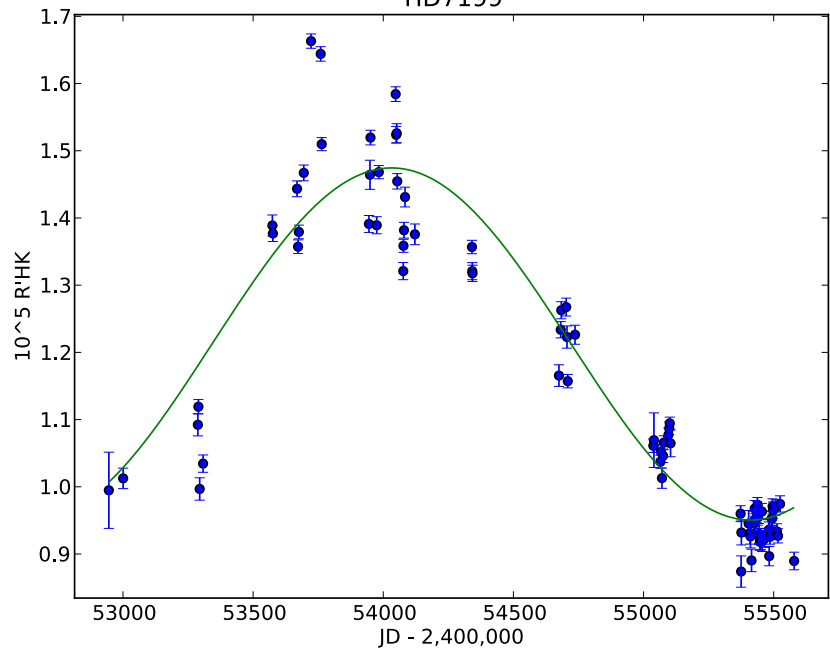
HD10700



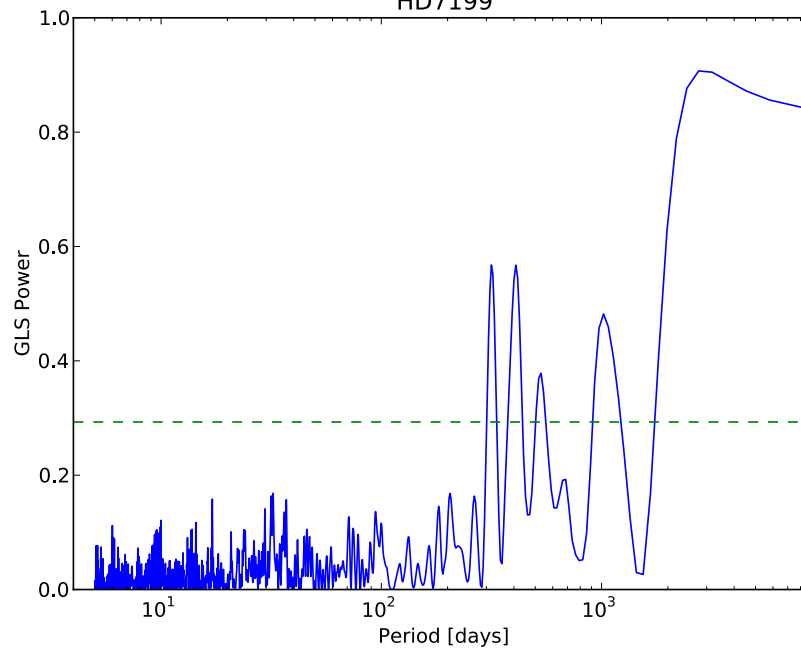
HD10700



HD7199



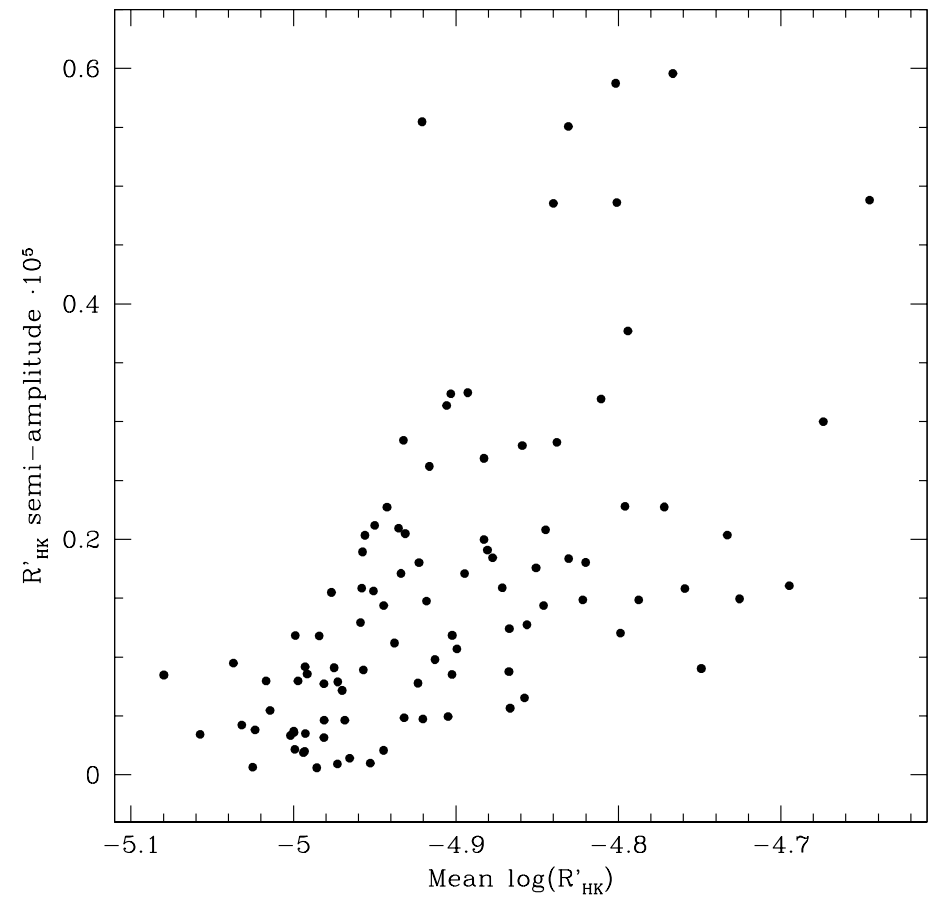
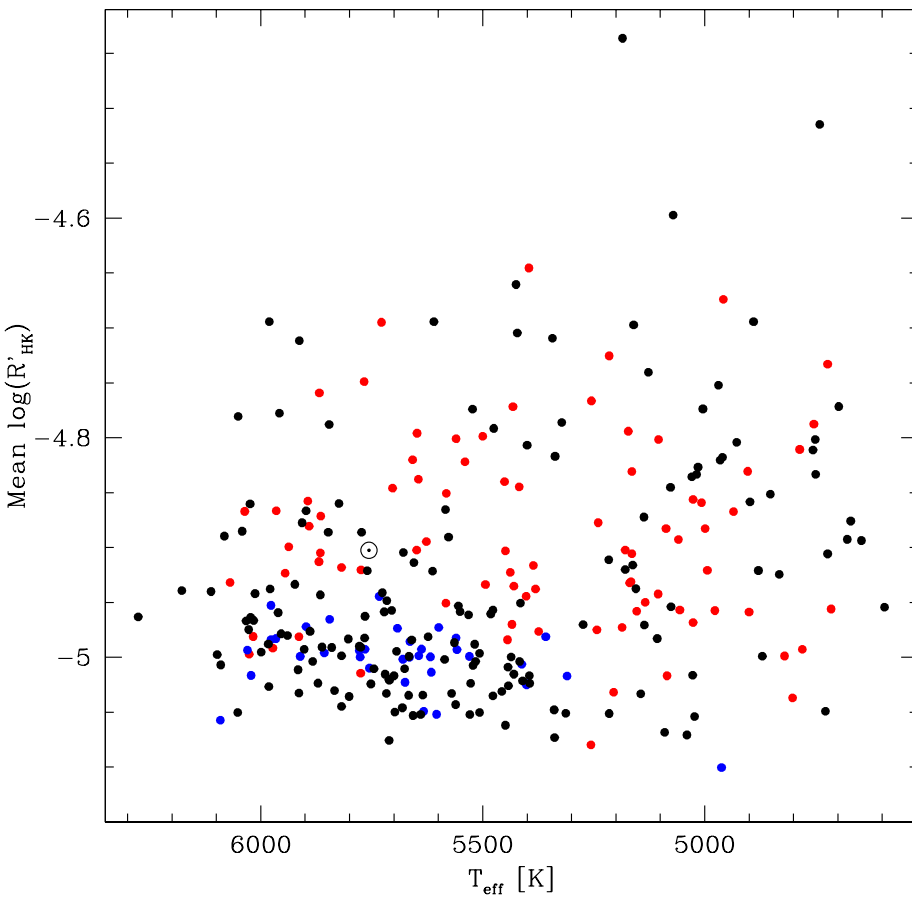
HD7199



Global results on magnetic cycles

- 284 main-sequence stars
 - 83 stars with cycle ($A > 0.04$)
 - 36 stars with no cycle ($A < 0.04$)
 - 165 stars with undetermined status
- 61 ± 8 % of older solar-type stars show a magnetic cycle, while 39 ± 8 % do not.

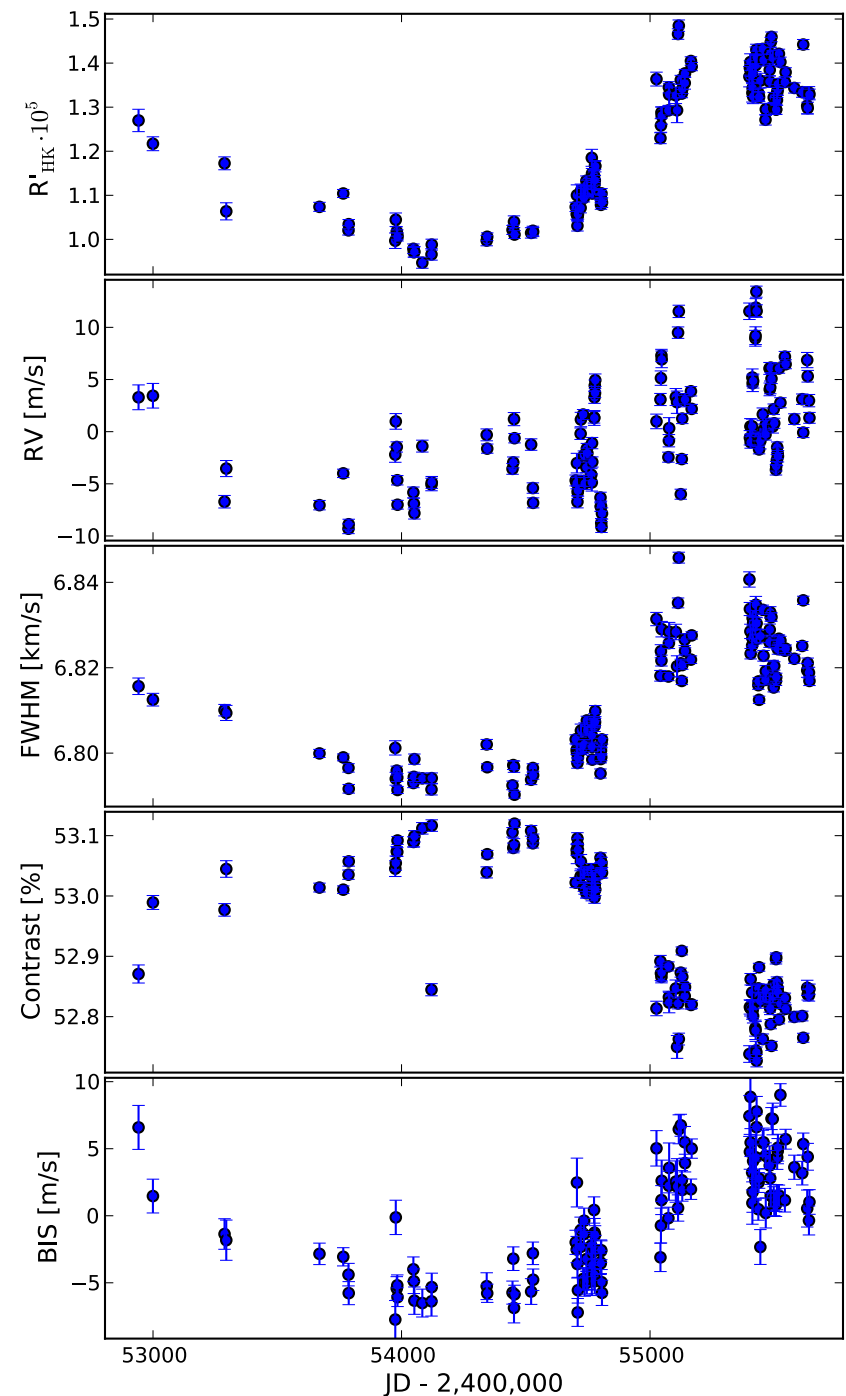
Global results on magnetic cycles



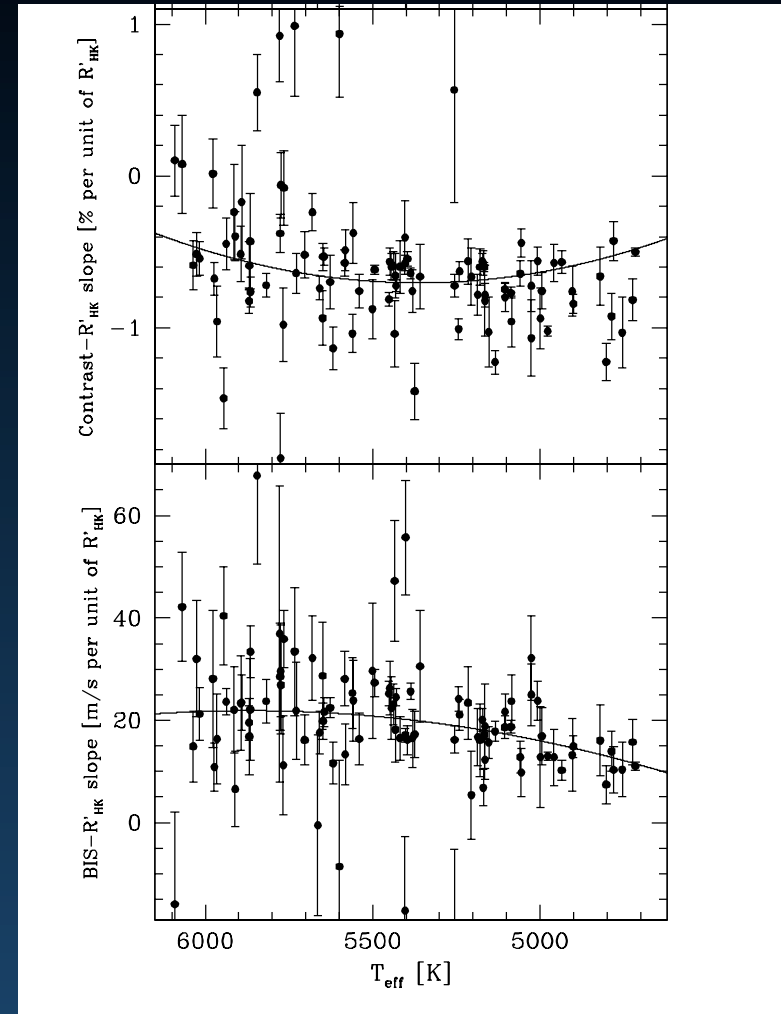
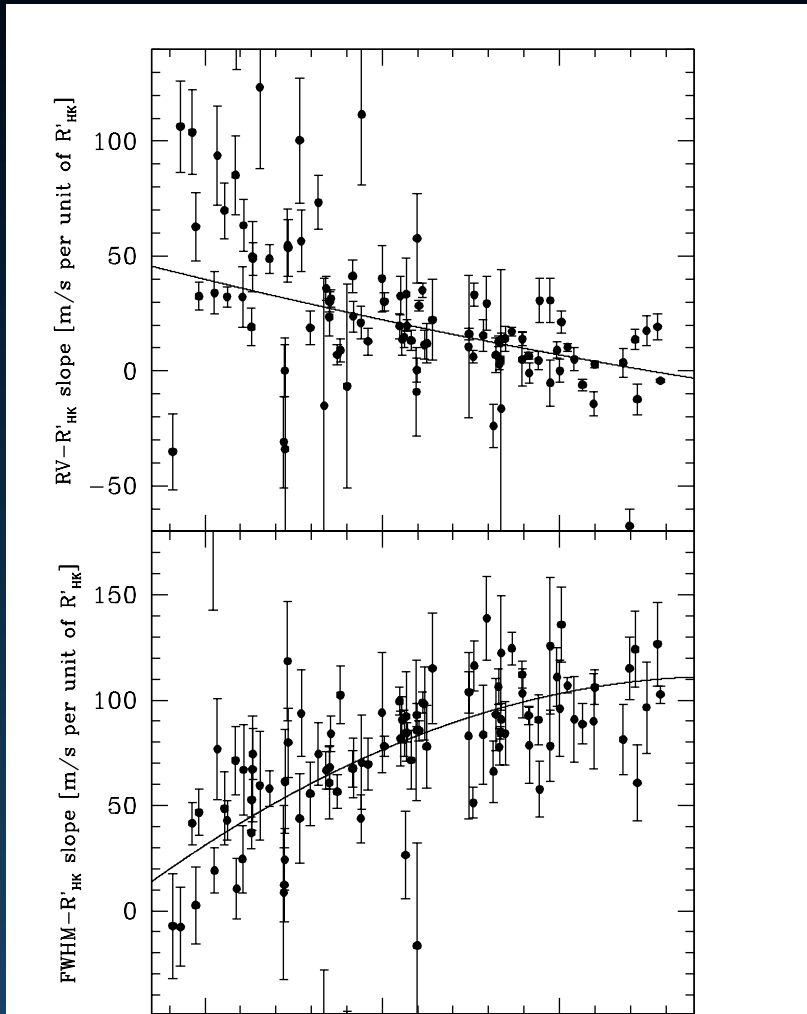
61 ± 8 % of older solar-type stars show a magnetic cycle, while 39 ± 8 % do not.

Impact of magnetic cycles on RV and CCF parameters

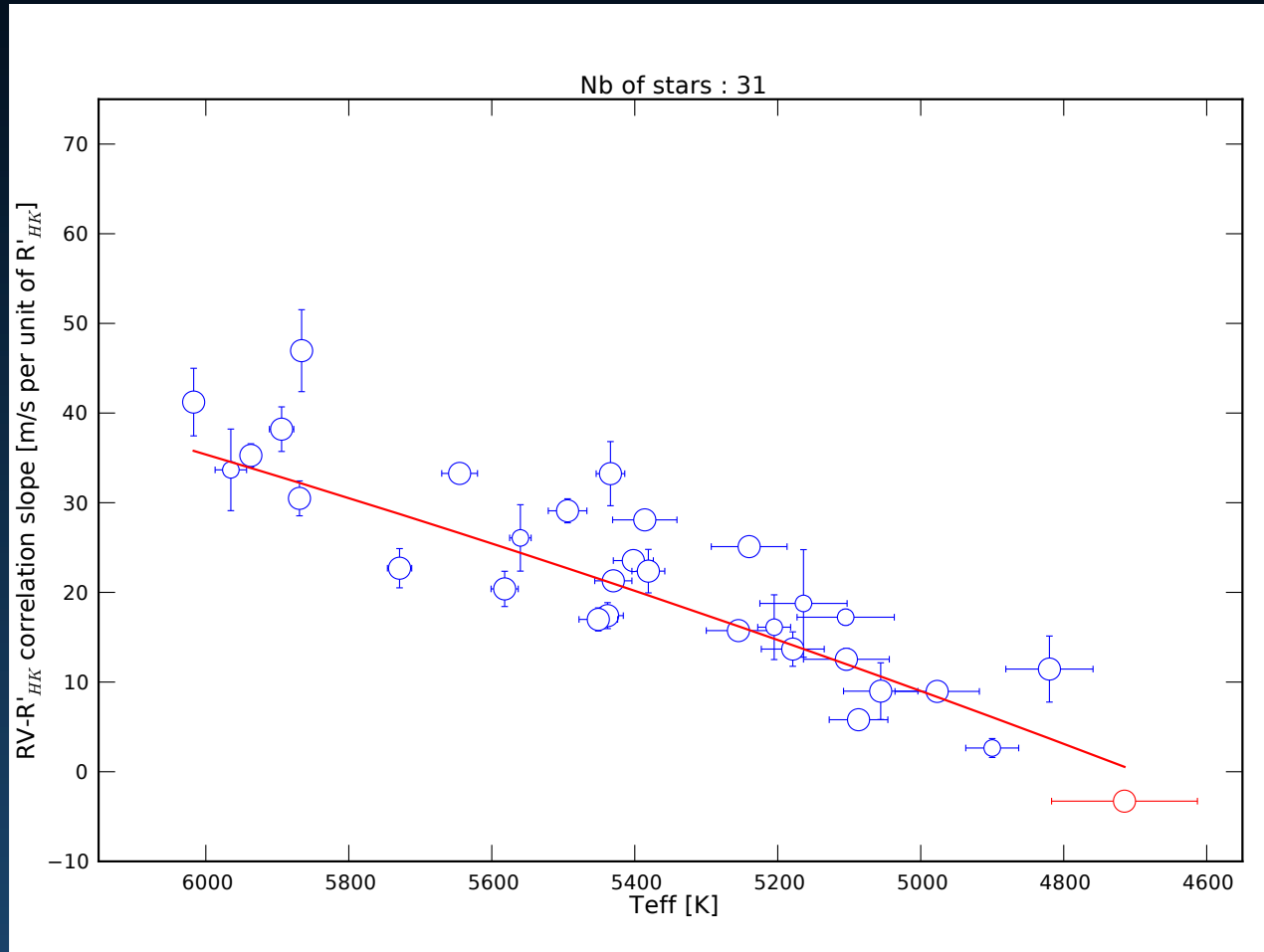
Example: HD 21693 (which by the way has two short-period planets, Mayor et al. 2011)



Sensitivity of CCF parameters to R'_{HK} variations



Sensitivity of radial velocities to R'_{HK} variations



Sensitivity of CCF parameters to R'_{HK} variations

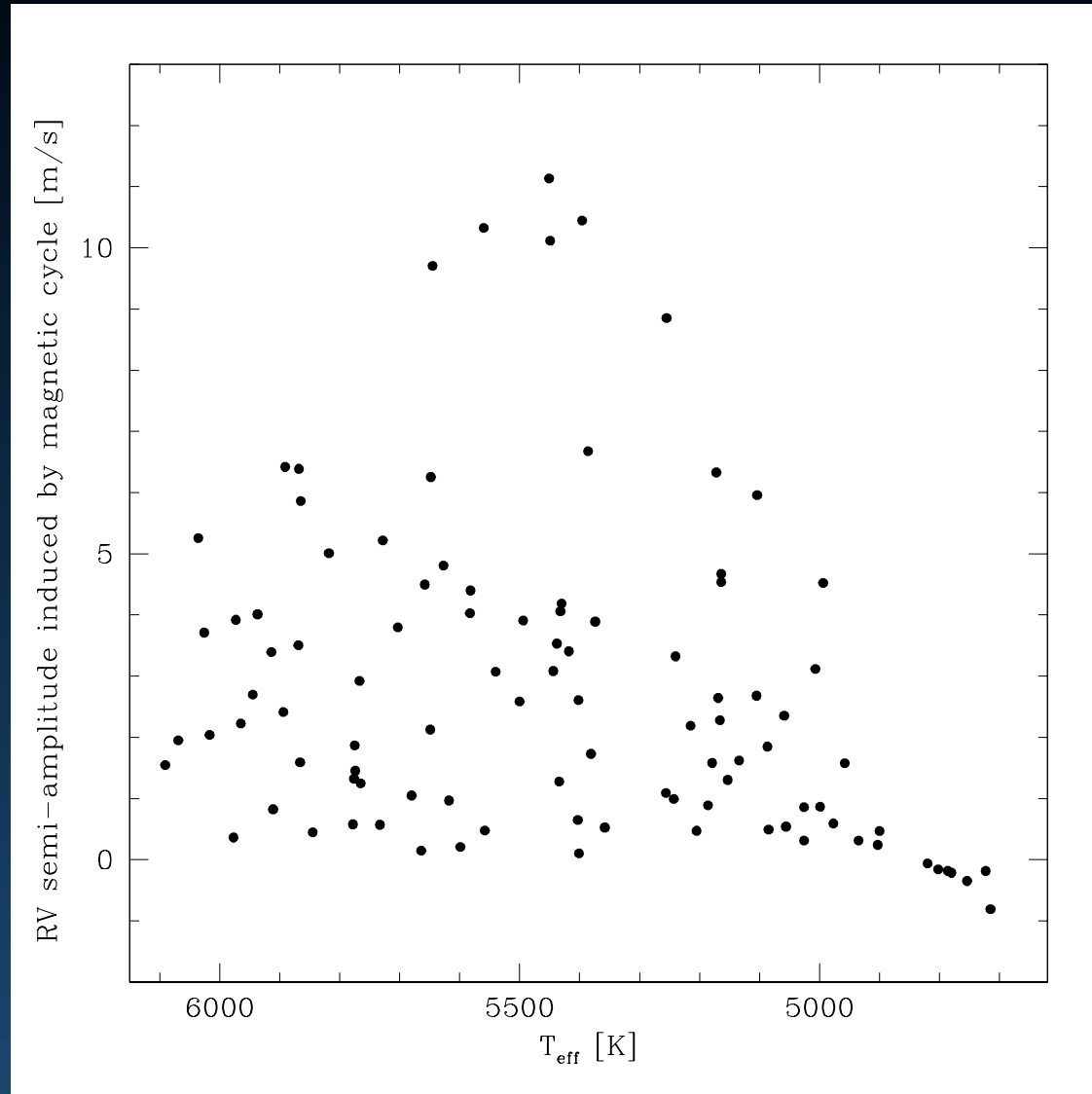
$$\begin{aligned} C_{\text{RV}} = & (19.04 \pm 0.62) + (3.20 \pm 0.20) \cdot 10^{-2} \cdot \tilde{T}_{\text{eff}} \\ & + (4.5 \pm 3.1) \cdot 10^{-6} \cdot \tilde{T}_{\text{eff}}^2 + (17.8 \pm 3.1) \cdot [\text{Fe}/\text{H}] \\ & + (24.3 \pm 5.2) \cdot [\text{Fe}/\text{H}]^2 \end{aligned} \quad (9)$$

$$\begin{aligned} C_{\text{FW}} = & (83.2 \pm 1.4) - (6.46 \pm 0.36) \cdot 10^{-2} \cdot \tilde{T}_{\text{eff}} \\ & - (3.71 \pm 0.79) \cdot 10^{-5} \cdot \tilde{T}_{\text{eff}}^2 + (33.7 \pm 6.8) \cdot [\text{Fe}/\text{H}] \\ & + (38 \pm 17) \cdot [\text{Fe}/\text{H}]^2 \end{aligned} \quad (10)$$

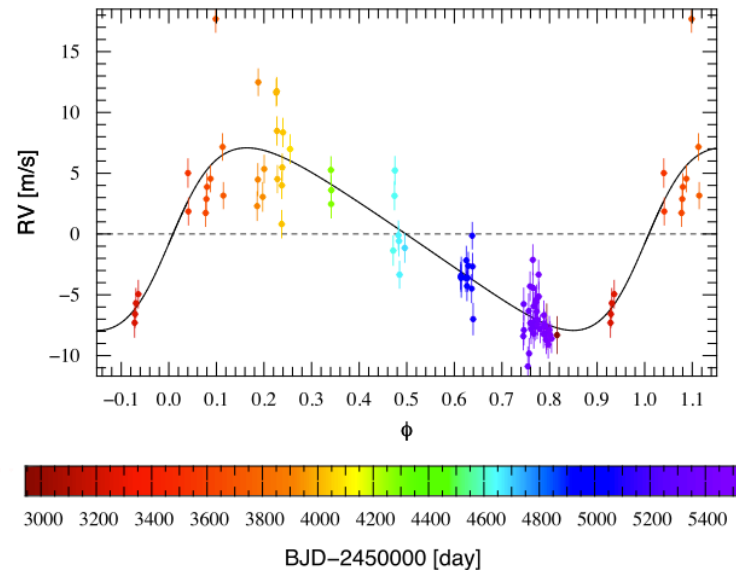
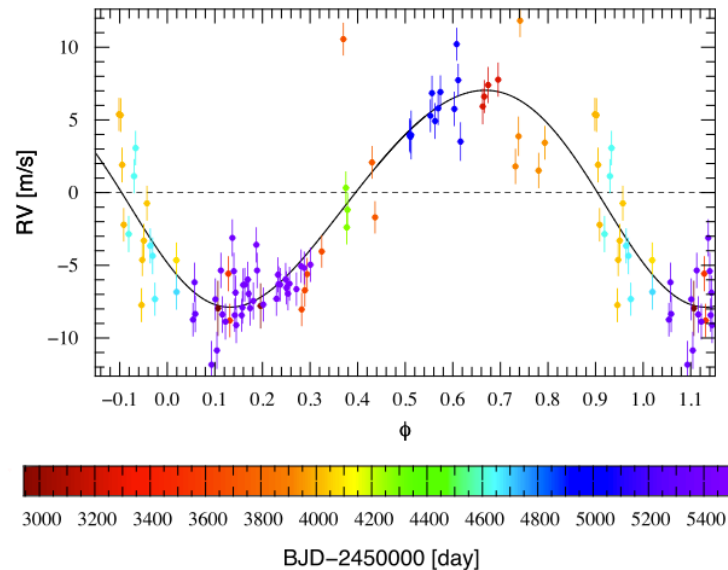
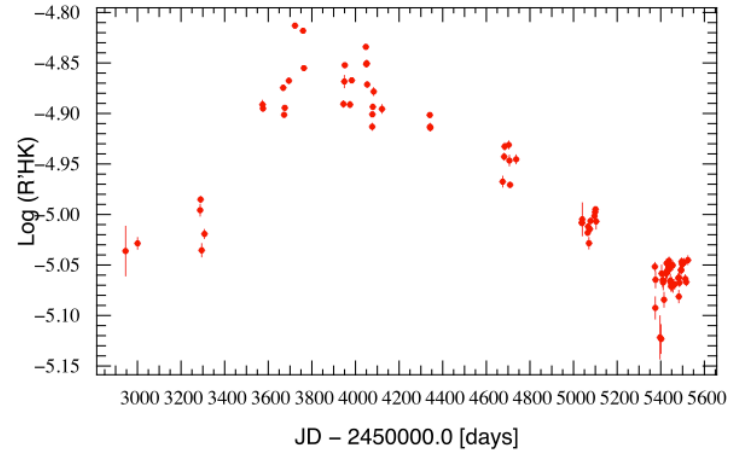
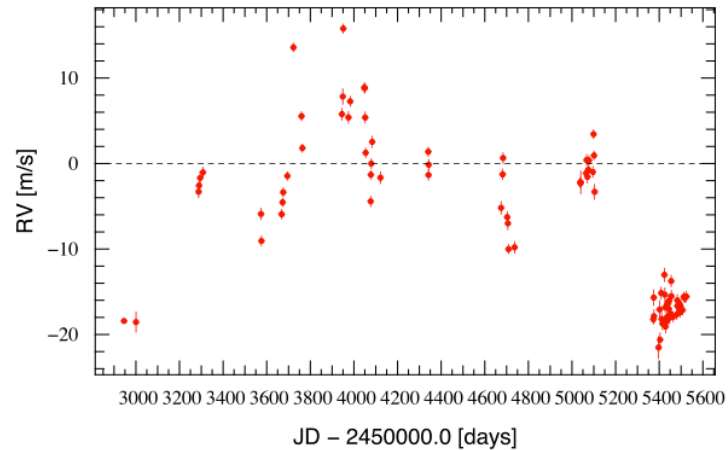
$$\begin{aligned} C_{\text{CO}} = & (-0.703 \pm 0.012) + (3.9 \pm 3.4) \cdot 10^{-5} \cdot \tilde{T}_{\text{eff}} \\ & + (5.28 \pm 0.67) \cdot 10^{-7} \cdot \tilde{T}_{\text{eff}}^2 + (0.333 \pm 0.055) \cdot [\text{Fe}/\text{H}] \\ & - (0.15 \pm 0.13) \cdot [\text{Fe}/\text{H}]^2 \end{aligned} \quad (11)$$

$$\begin{aligned} C_{\text{BIS}} = & (20.18 \pm 0.44) + (7.3 \pm 1.3) \cdot 10^{-3} \cdot \tilde{T}_{\text{eff}} \\ & - (7.8 \pm 2.4) \cdot 10^{-6} \cdot \tilde{T}_{\text{eff}}^2 + (8.2 \pm 2.3) \cdot [\text{Fe}/\text{H}] \\ & + (13.7 \pm 5.1) \cdot [\text{Fe}/\text{H}]^2 \end{aligned} \quad (12)$$

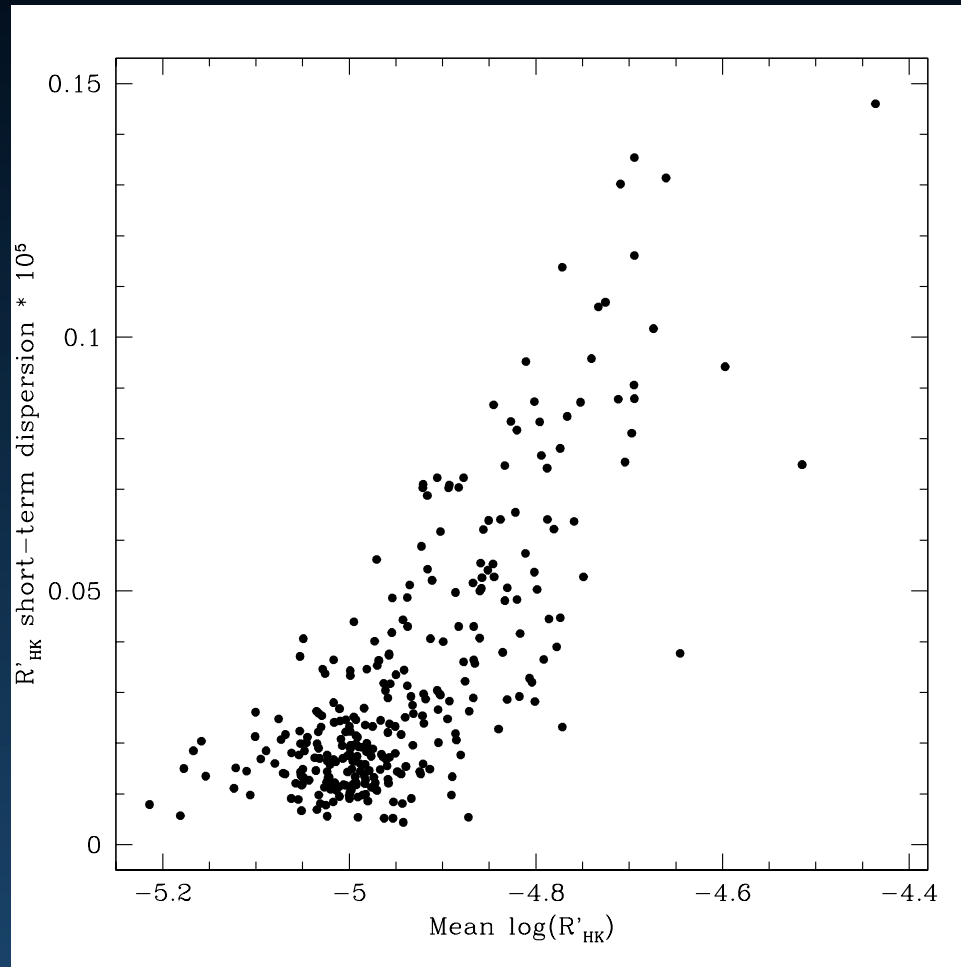
Effective impact on radial velocities



Disentangling planets and magnetic cycles: the example of HD 7199 (Dumusque et al. 2011)



Short-term R'_{HK} variations: rotational modulations



Short-term R'_{HK} variations: rotational modulations

*See poster 20.04 by X. Dumusque on
Alpha Cen B!*

*« All included » package:
binarity, magnetic cycle, rotational
modulations, barycentric correction, and...
low-mass planets around Alpha Cen B ?!*