## gravitySimulator: Beyond the Million-Body Problem

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One of the most cpu-intensive calculations in astrophysics is the gravitational N-body problem. The N-body problem is particularly challenging in the case of galactic nuclei, because of the steeply-rising stellar density profile and the presence of single or multiple supermassive black holes. Not only must the particle advancement be very accurate, but the value of *N* must be chosen large enough that two-body scattering does not artificially repopulate the loss cone of the central object on time scales shorter than the orbital period. This requirement imposes a minimum value for N, which for typical applications is of order 106 or greater

The state-of-the-art way to deal with such large particle numbers is via the GRAPE special-purpose computers. But the finite on-board memory of the GRAPEs limits the number of particles that can be handled. This limitation can be overcome by linking multiple GRAPE boards into a cluster. Such a cluster ("gravitySimulator") has recently become operational at the Rochester Institute of Technology

In this poster, we compare different ways of implementing a parallel *N*-body code on the GRAPE cluster and test their



g 1: A single GRAPE-6A card fits into the PCI-slot of a ommon PC and can accelerate the force calculation in an *N*-ody simulation by a factor of 100. The peak-performance is 25Gflops but the card's memory can only hold 128k particles.



Fig 2: The RIT GRAPE cluster "gravitySimulator" is in operation since February 2005. Here are some of the technical operati details:

- 1 head node and 32 computing nodes • dual 3GHz Xeon processors with 2Gbyte of memory
  • 32 GRAPE-6A cards
- 14 Tbyte RAID arrayfast low-latency InfiniBand interconnect
- 10Gbps theoretical peak-performance is 4Gflops N up to 4x10<sup>6</sup>
  Cost: \$0.5x10<sup>6</sup>

- Funding: NSF/NASA/RIT Next largest: 24 nodes (University of Tokyo) • A similar 32-node cluster will soon be operational in Heidelberg, Germany



Berczik P., Merrit D., Spurzem R., 2005, astro-ph/0507260 Dorband E.N., Hemsendorf M., Merritt D., 2003, JCP, 185, 484 Gualandris A., Portegies-Zwart S., Tirado-Ramos A., 2004, IEEE Harfst S., Merritt D., Gualandris A., Berczik P., 2005, in preparation

For further information and updates about on-going projects including a new visualization tool visit the website of the GRAPE cluster project at http://www.grapecluster.rit.edu/

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nen (γ=1.5) 104 10<sup>5</sup> 106 Number of Particles Ν g 4: Timing results from simulations using two different

Dehnen models. The wallclock time is measured for one full Dennen models. The wallclock time is measured for one full time step  $\lambda$ cell. Generally, the more concentrated model ( $r \in 1.5$ ) take longer to compute. For larger particle numbers (roughly  $N > 10^{\circ}$ ) the wallclock time scales with  $N^{\circ}$  as expected and using additional processors decreases computation time. Below 10<sup>4</sup> particles computation time increases for more processors due to the overchand in computation time. to the overhead in communication





Fig 5: Efficiency – defined as computation time needed on one processor divided by *p* times the time needed on *p* processors – for different particle numbers. Unit (perfect) efficiency corresponds to zero communication and latency losses. For low particle numbers, the force calculation is very Tast and communication dominates the run time. The code runs most efficiently if the number of particles on each node is close to the maximum (128k) permitted by the GRAPE memory.